

Component Manual for the Xray-Tracing Package McXtrace, version 1.7.1

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The software package McXtrace is a tool for carrying out highly complex Monte Carlo ray-tracing simulations of X-ray beamlines to high precision. The simulations can compute all aspects of the performance of instruments and can thus be used to optimize the use of existing equipment, design new instrumentation, and carry out virtual experiments for e.g. training, experimental planning or data analysis. McXtrace is based is based on a unique design, inhereted from its sister McStas, where an automatic compilation process translates high-level textual instrument descriptions into efficient ANSI-C code. This design makes it simple to set up typical simulations and also gives essentially unlimited freedom to handle more unusual cases. This report constitutes the component manual for McXtrace, and, together with the manual for the McXtrace system, it contains full documentation of all aspects of the program. It covers a description of all official components of the McXtrace package with some theoretical background. Selected test instruments and representative McXtrace simulations performed with these instruments are described in the User Manual.

This report documents the components for McXtrace version 1.7.1, released Dec. 2022.

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Contents

	Pref	face and acknowledgements	8
1.	Abo	ut the component library	9
	1.1.	Authorship	9
	1.2.	Symbols for x-ray scattering and simulation	9
		Component coordinate system	10
	1.4.	About data files	10
	1.5.	Component source code	11
	1.6.	Documentation	11
	1.7.	Disclaimer, bugs	11
2.	Mor	nte Carlo Techniques and simulation strategy	12
	2.1.	X-ray simulations	12
		2.1.1. Monte Carlo ray tracing simulations	13
	2.2.	The x-ray weight	13
		2.2.1. Statistical errors of non-integer counts	14
	2.3.	Weight factor transformations during a Monte Carlo choice	15
		2.3.1. Direction focusing	15
	2.4.	Stratified sampling	16
		Accuracy of Monte Carlo simulations	17
3.	Sou	rce components	18
		3.0.1. Photon flux and Brilliance	18
	3.1.	Source_pt: A mathematical point emitting photons with a spectrum either	
		uniform, gaussian or generated from a datafile	20
	3.2.	Source_flat: A flat surface emitting photons with a spectrum either uni-	
		form, gaussian or generated from a datafile	21
	3.3.	Source_div: A continuous source with specified divergence	21
	3.4.	Source_gaussian: the model has a gaussian distribution of intensity	22
	3.5.	Source_lab: X-ray tube laboratory source	23
	3.6.	Other sources components: virtual sources (event files)	27
4.	Bea	m optical components: Arms, slits, filters etc.	28
	4.1.	Arm: The generic component	28
		Slit: A beam defining diaphragm	
		Slit_N: multiple slits	
		Beamstop: A photon absorbing area	

	4.5.	Filter: A general absorption filter model	
	4.6.	Chopper_simple: An ideal chopper	
5.	Refr	active optical components: lenses	33
	5.1.	Lens_simple: Thin lens approximation	33
	5.2.	Lens_parab: Thick parabolic CRL	33
	5.3.	Lens_parab_Cyl: Thick 1D-parabolic CRL	34
	5.4.	Lens_parab_rough: Thick parabolic CRL including roughness-model	34
	5.5.	$Lens_parab_Cyl_rough:\ Thick\ 1D-parabolic\ CRL\ including\ roughness-model$	35
	5.6.	Lens_Kinoform: refractice kinoform lens	35
	5.7.	Lens_elliptical:	36
6.	Refle	ective optical components: mirrors	37
	6.1.	Mirror_curved: Cylindrically curved mirror	37
	6.2.	Mirror_parabolic: Mirror with a parabolic curvature profile	37
	6.3.	Mirror_elliptic: Mirror with a elliptic curvature profile	38
	6.4.	Multilayer_elliptic: Elliptically curved mirror coated with a multilayer	38
		6.4.1. Definition of the reference frames	39
		6.4.2. Algorithm	39
	6.5.	Reflection of the ray in the mirror	40
		6.5.1. Mirror reflectivity	41
	6.6.	TwinKB_ML: Side-by-side Kirkpatrick-Baez mirror pair	41
7.	Sam	•	43
		7.0.1. Scattering notation	43
		7.0.2. Weight transformation in samples; focusing	44
		7.0.3. Future development of sample components	45
	7.1.	Absorption_sample: An absorption phantom	45
	7.2.	Saxs_spheres: A model of dilute hard spheres in solution for SAXS-use	47
	7.3.	PowderN: A general powder sample	48 49
		7.3.2. Geometry, physical properties, concentricity	
		7.3.3. Powder scattering	50
		7.3.4. Algorithm	52
	7.4.	Perfect_crystal: A Darwin-width domniated single crystal model	$\frac{52}{54}$
	7.5.	Single_crystal: The single crystal component	55
	7.6.	Molecule_2state: Excitable time-dependent sample model	57
8.	Mon	itors and detectors	58
J.	8.1.	Monitor: Simple intensity monitor	59
	8.2.	E_monitor: The energy-sensitive monitor	59
		L_monitor: The wavelength sensitive monitor	59
		PSD monitor: The PSD monitor	60

	8.5.	PSD_monitor_coh: The coherent PSD monitor	61
	8.6.	PSD_monitor_4PI: A 4π sr spherical monitor	61
	8.7.	EPSD_monitor: Energy-selective PSD monitor	62
	8.8.	W_psd_monitor: A power vs. position monitor	62
	8.9.	Monitor_nD: A general Monitor for 0D/1D/2D records	64
		8.9.1. The Monitor_nD geometry	65
		8.9.2. The photon parameters that can be monitored	66
		8.9.3. Important options	66
		8.9.4. The output files	67
		8.9.5. Monitor equivalences	67
		8.9.6. Usage examples	68
		8.9.7. Monitoring user variables	69
	8.10.	. PreMonitor_nD: Store photon rays for possible later detection	70
9.	Spec	cial-purpose components	72
		Progress_bar: Dynamic information output	73
	9.2.	Virtual_output: Saving the first part of a split simulation	73
	9.3.	Virtual_input: Starting the second part of a split simulation	73
	9.4.	Shadow_input: Reading input from Shadow	74
	9.5.	Shadow_output: Saving the photon rays for use with SHADOW	74
Α.		aries and conversion constants	75
	A.1.	Run-time calls and functions (mcxtrace-r)	75
		A.1.1. Photon propagation	75
		A.1.2. Coordinate and component variable retrieval	76
		A.1.3. Coordinate transformations	77
		A.1.4. Mathematical routines	78
		A.1.5. Output from detectors	78
		A.1.6. Ray-geometry intersections	79
		A.1.7. Random numbers	79
		Reading a data file into a vector/matrix (Table input, read_table-lib) .	80
	A.3.	Constants for unit conversion etc	83
В.	The	McXtrace terminology	85
	Bibli	iography	86

Preface and acknowledgements

This document contains information on the x-ray scattering components which are the building blocks for defining instruments in the Monte Carlo Xray-tracing program McX-trace version 1.7.1. The initial release in June 2011 of version 1.0 was presented in Ref. [LN99]. The reader of this document is not supposed to have specific knowledge of xray scattering, but some basic understanding of physics is helpful in understanding the theoretical background for the component functionality. For details about setting up and running simulations, we refer to the McXtrace system manual [Knu+14]. We assume familiarity with the use of the C programming language.

We would like to explicitly thank all the partners in this project:

- The European Synchrotron Radiation Facility (ESRF), Grenoble, France
- SAXSLAB Aps., Lundtofte, Denmark
- Risø DTU, Roskilde, Denmark
- Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark.
- Faculty for Life Sciences, University of Copenhagen, Copenhagen, Denmark.

As McXtrace has inherited much of its functionality from its sister McStas we take the oppurtunity to thank Dir. Kurt N. Clausen, PSI, for his continuous support to McStas and for having initiated the project. Continuous support to McStas has also come from Prof. Robert McGreevy, ISIS.

We have further benefited from discussions with many other people in the scattering community, too numerous to mention here.

The users who contributed components to this manual are acknowledged as authors of the individual components. We encourage other users to contribute components with manual entries for inclusion in future versions of McXtrace.

In case of errors, questions, or suggestions, do not hesitate to contact the user/developer community by writing to the user mailing list mcxtrace-users@mcxtrace.órg or consult the McXtrace home page [Mcx]. A special developement website (shared with the sister project McStas) complete with bug/request reporting service is available [Tra].

We would like to kindly thank all McXtrace component contributors. This is the way we improve the software alltogether.

If you appreciate this software, please subscribe to the mcxtrace-users@mcxtrace.org email list, send us a smiley message, and contribute to the package.

We also encourage you to refer to this software when publishing results, with the following citation:

• E. B. Knudsen, et.al, J. Applied Cryst. 46, pp 679, 2013.

1. About the component library

This McXtrace Component Manual consists of the following major parts:

- An introduction to the use of Monte Carlo methods in McXtrace.
- A thorough description of system components, with one chapter per major category: Sources, optics, monochromators, samples, monitors, and other components.
- The McXtrace library functions and definitions that aid in the writing of simulations and components in Appendix A.
- An explanation of the McXtrace terminology in Appendix B.

Additionally, you may refer to the list of example instruments from the library in the McXtrace User Manual.

1.1. Authorship

The component library is maintained by the McXtrace system group. A number of basic components "belongs" the McXtrace system, and are supported and tested by the McXtrace team.

Other components are contributed by specific authors, who are listed in the code for each component they contribute as well as in this manual. McXtrace users are encouraged to send their contributions to us for inclusion in future releases.

1.2. Symbols for x-ray scattering and simulation

In the description of the theory behind the component functionality we will use the usual symbols \mathbf{r} for the position (x, y, z) of the particle (unit m), and \mathbf{k} for the particle wave vector (k_x, k_y, k_z) (unit $\mathring{\mathbf{A}}^{-1}$). The wavelength is the reciprocal wave vector, $\lambda = 2\pi/k$. By convention energy is usually given i keV and may be calculated from the wavelength by: $\lambda = \frac{12.398}{E}$

In general, vectors are denoted by boldface symbols.

Subscripts "i" and "f" denotes "initial" and "final", respectively, and are used in connection with the photon state before and after an interaction with the component in question.

MCXTRACE/data	Description
*.lau	Laue pattern file, as issued from Crystallographica. For
	use with Single_crystal and PowderN. Data: [h k l Mult.
	d-space 2Theta F-squared]
*.laz	Powder pattern file, as obtained from Lazy/ICSD. For
	use with PowderN.
*.txt	General text file, containing ascii data. Currently used
	for elemental data extracted from NIST[Nis].

Table 1.1.: Data files of the McXtrace library.

1.3. Component coordinate system

All mentioning of component geometry refer to the local coordinate system of the individual component. The axis convention is so that the z axis is along the photon propagation axis, the y axis is vertical up, and the x axis points left when looking along the z-axis, completing a right-handed coordinate system. Most components 'position' (as specified in the instrument description with the AT keyword) corresponds to their volume centre. The mirrors are an important counterexmaple. In this case the 'position' is the centre of the mirror surface

Components are not necessarily designed to overlap. This may lead to loss of rays. Warnings will be issued during simulation if sections of the instrument are not reached by any xrays, or if a significant number of xrays are removed. This is usually the sign of either overlapping components or low intensity.

1.4. About data files

Some components require external data files, e.g. lattice crystallographic definitions for Laue and powder pattern diffraction, absorption and reflectivity files, etc.

Such files distributed with McXtrace are located in the data sub-directory of the MCXTRACE library. Components that make use of the McXtrace file system, including the read-table library (see section A.2) may access all McXtrace data files without making local copies. Of course, you are welcome to define your own data files, and eventually contribute to McXtrace if you find them useful.

File extensions are not interpreted by McXtrace but help in identifying relevant files per application. Table 1.1 lists the current default file extensions.

McXtrace itself generates both simulation and monitor data files, the structure of which is explained in the User Manual (see end of chapter 'Running McXtrace').

1.5. Component source code

Source code for all components may be found in the MCXTRACE library subdirectory of the McXtrace installation; the default is /usr/local/lib/mcxtrace/ on Unix-like systems and C:\mcxtrace\lib on Windows systems, but it may be changed using the MCXTRACE environment variable.

In case users only require to add new features, preserving the existing features of a component, using the EXTEND keyword in the instrument description file is recommended. For larger modification of a component, it is advised to make a copy of the component file into the working directory. A component file in the local directory will in McXtrace takes precedence over a library component of the same name.

1.6. Documentation

As a complement to this Component Manual, we encourage users to use the mxdoc front-end which enables to display both the catalog of the McXtrace library, e.g using:

mxdoc

as well as the documentation of specific components, e.g with:

```
mxdoc --text name
mxdoc file.comp
```

The first line will search for all components matching the *name*, and display their help section as text. For instance, mxdoc .laz will list all available Lazy data files, whereas mxdoc --text Monitor will list most Monitors. The second example will display the help corresponding to the *file.comp* component, using your BROWSER setting, or as text if unset. The --help option will display the command help, as usual.

An overview of the component library is also given at the McXtrace home page [Mcx] and in the User Manual [Knu+14].

1.7. Disclaimer, bugs

We would like to emphasize that the usage of both the McXtrace software, as well as its components are the responsability of the users. Indeed, obtaining accurate and reliable results requires a substantial work when writing instrument descriptions. This also means that users should read carefully both the documentation from the manuals [Knu+14] and from the component itself (using mcdoc comp) before reporting errors. Most anomalous results often originate from a wrong usage of some part of the package.

Anyway, if you find that either the documentation is not clear, or the behavior of the simulation is undoubtedly anomalous, you should report this to us at mcxtrace-users@mcxtrace.org and/or refer to our special bug/request reporting service [Mcz].

2. Monte Carlo Techniques and simulation strategy

This chapter explains the simulation strategy and the Monte Carlo techniques used in McXtrace. We first explain the concept of the x-ray weight factor, and discuss the statistical errors in dealing with sums of x-ray weights. Secondly, we give an expression for how the weight factor transforms under a Monte Carlo choice and specialize this to the concept of direction focusing. Finally, we present a way of generating random numbers with arbitrary distributions. More details are available in the Appendix concerning random numbers in the User manual.

2.1. X-ray simulations

X-ray scattering beamlines are built as a series of optical elements. Each of these elements modifies the beam characteristics (e.g. divergence, wavelength spread, spatial and temporal distributions) in a way which, for simple x-ray beam configurations, may be modelled with analytical methods.

However, real x-ray beamlines consist of a large number of optical elements, and this brings additional complexity by introducing strong correlations between x-ray beam parameters like divergence and position - which is the basis of the acceptance diagram method - but also wavelength and time. The usual analytical methods, such as phase-space theory, then reach their limit of validity in the description of the resulting effects.

In order to cope with this difficulty, Monte Carlo (MC) methods (for a general review, see Ref. [Jam80]) may be applied to the simulation of x-ray beamlines. The use of probability is commonplace in the description of microscopic physical processes. Integrating these events (absorption, scattering, reflection, ...) over the x-ray trajectories results in an estimation of measurable quantities characterizing the beamline. Moreover, using variance reduction (importance sampling) where possible, reduces the computation time and gives better accuracy.

Implementations of the MC method for X-ray beamlines already exist, most notable is probably SHADOW[WCC94], originally developed by the late Franco Cerrina and coworkers, now developed further by M. Sanchez Del Rio at the ESRF[Rio+11][Sha]. Other implementations of the same concept are RAY[Sch08] from BESSY and Xtrace[Bau+07]. hosted at ANKA

2.1.1. Monte Carlo ray tracing simulations

Mathematically, the Monte-Carlo method is an application of the law of large numbers [Jam80; GRR92]. Let f(u) be a finite continuous integrable function of parameter u for which an integral estimate is desirable. The discrete statistical mean value of f (computed as a series) in the uniformly sampled interval a < u < b converges to the mathematical mean value of f over the same interval.

$$\lim_{n \to \infty} \frac{1}{n} \sum_{i=1, a < u_i < b}^{n} f(u_i) = \frac{1}{b-a} \int_a^b f(u) du$$
 (2.1)

In the case were the u_i values are regularly sampled, we come to the well known midpoint integration rule. In the case were the u_i values are randomly (but uniformly) sampled, this is the Monte-Carlo integration technique. As random generators are not perfect, we rather talk about quasi-Monte-Carlo technique. We encourage the reader to refer to James [Jam80] for a detailed review on the Monte-Carlo method.

2.2. The x-ray weight

A totally realistic semi-classical simulation will require that each x-ray is at any time either present or lost. On many beamlines the sheer abundance of x-ray photons makes it impractical to trace each and every photon from the source. This is particularly the case at XFELs. Additionally, only a very small fraction of the initial x-rays will ever be detected, and simulations of this kind will therefore waste much time in dealing with x-rays that never hit the detector.

A way of dealing with these issues and speed up calculations is to introduce a x-ray "weight factor" for each simulated ray and to adjust this weight according to the path of the ray. If e.g. the reflectivity of a certain optical component is 10%, and only reflected x-rays ray are considered later in the simulations, the x-ray weight will be multiplied by 0.10 when passing this component, but every x-ray is allowed to reflect in the component. In contrast, the totally realistic simulation of the component would require on average ten incoming x-rays for each reflected one.

Let the initial x-ray weight be p_0 and let us denote the weight multiplication factor in the j'th component by π_j . The resulting weight factor for the x-ray ray after passage of the whole beamline becomes the product of all contributions

$$p = p_n = p_0 \prod_{j=1}^n \pi_j. (2.2)$$

Each adjustement factor should be $0 < \pi_j < 1$, except in special circumstances, so that total flux can only decrease through the simulation. For convenience, the value of p is updated (within each component) during the simulation.

Simulation by weight adjustment is performed whenever possible. This includes

• Transmission through filters and windows.

- Reflection from monochromator (and analyser) crystals with finite reflectivity and mosaicity.
- Reflections from mirrors.
- Passage of a continuous beam through a chopper.
- Scattering from all types of samples.

2.2.1. Statistical errors of non-integer counts

In a typical simulation, the result will consist of a count of x-ray histories ("rays") with different weights. The sum of these weights is an estimate of the mean number of x-rays hitting the monitor (or detector) per second in a "real" experiment. One may write the counting result as

$$I = \sum_{i} p_i = N\overline{p},\tag{2.3}$$

where N is the number of rays hitting the detector and the horizontal bar denotes averaging. By performing the weight transformations, the (statistical) mean value of I is unchanged. However, N will in general be enhanced, and this will improve the accuracy of the simulation.

To give an estimate of the statistical error, we proceed as follows: Let us first for simplicity assume that all the counted x-ray weights are almost equal, $p_i \approx \overline{p}$, and that we observe a large number of x-rays, $N \geq 10$. Then N almost follows a normal distribution with the uncertainty $\sigma(N) = \sqrt{N}^{-1}$. Hence, the statistical uncertainty of the observed intensity becomes

$$\sigma(I) = \sqrt{N}\overline{p} = I/\sqrt{N},\tag{2.4}$$

as is used in real x-ray experiments (where $\bar{p} \equiv 1$). For a better approximation we return to Eq. (2.3). Allowing variations in both N and \bar{p} , we calculate the variance of the resulting intensity, assuming that the two variables are independent:

$$\sigma^2(I) = \sigma^2(N)\overline{p}^2 + N^2\sigma^2(\overline{p}). \tag{2.5}$$

Assuming as before that N follows a normal distribution, we reach $\sigma^2(N)\overline{p}^2 = N\overline{p}^2$. Further, assuming that the individual weights, p_i , follow a Gaussian distribution (which in some cases is far from the truth) we have $N^2\sigma^2(\overline{p}) = \sigma^2(\sum_i p_i) = N\sigma^2(p_i)$ and reach

$$\sigma^{2}(I) = N\left(\overline{p}^{2} + \sigma^{2}(p_{i})\right). \tag{2.6}$$

The statistical variance of the p_i 's is estimated by $\sigma^2(p_i) \approx (\sum_i p_i^2 - N\overline{p}^2)/(N-1)$. The resulting variance then reads

$$\sigma^2(I) = \frac{N}{N-1} \left(\sum_i p_i^2 - \overline{p}^2 \right). \tag{2.7}$$

¹This is not correct in a situation where the detector counts a large fraction of the x-rays in the simulation, but we will neglect that for now.

For almost any positive value of N, this is very well approximated by the simple expression

$$\sigma^2(I) \approx \sum_i p_i^2. \tag{2.8}$$

As a consistency check, we note that for all p_i equal, this reduces to eq. (2.4)

In order to compute the intensities and uncertainties, the detector components in McXtrace will keep track of $N = \sum_i p_i^0$, $I = \sum_i p_i^1$, and $M_2 = \sum_i p_i^2$.

2.3. Weight factor transformations during a Monte Carlo choice

When a Monte Carlo choice must be performed, e.g. when the initial energy and direction of the x-ray ray is decided at the source, it is important to adjust the x-ray weight so that the combined effect of x-ray weight change and Monte Carlo probability of making this particular choice equals the actual physical properties we like to model.

Let us follow up on the simple example of transmission. The probability of transmitting the real x-ray is P, but we make the Monte Carlo choice of transmitting the x-ray every time: $f_{\text{MC}} = 1$. This must be reflected on the choice of weight multiplier π_j given by the master equation In the "real" semi-classical world, there is a distribution (probability density) for the x-rays in the six dimensional (energy, direction, position) space of $\Pi(E, \mathbf{\Omega}, \mathbf{r}) = dP/(dEd\mathbf{\Omega}d^3\mathbf{r})$ depending upon the source type and its parameters (such as gap, period, field strength etc. for an undulator). In the Monte Carlo simulations, the six coordinates are for efficiency reasons in general picked from another distribution: $f_{\text{MC}}(E, \mathbf{\Omega}, \mathbf{r}) \neq \Pi(E, \mathbf{\Omega}, \mathbf{r})$, since one would e.g. often generate only x-rays within a certain parameter interval. However, we must then require that the weights are adjusted by a factor π_j (in this case: j=1) so that

$$f_{\rm MC}\pi_i = P. \tag{2.9}$$

This probability rule is general, and holds also if, e.g., it is decided to transmit only half of the rays ($f_{\rm MC}=0.5$). An important different example is elastic scattering from a powder sample, where the Monte-Carlo choices are the particular powder line to scatter from, the scattering position within the sample and the final x-ray direction within the Debye-Scherrer cone.

2.3.1. Direction focusing

An important application of weight transformation is direction focusing. Assume that the sample scatters the x-rays in many directions. In general, only x-rays in some of these directions will stand any chance of being detected. These directions we call the *interesting directions*. The idea in focusing is to avoid wasting computation time on x-rays scattered in the other directions. This trick is an instance of what in Monte Carlo terminology is known as *importance sampling*.

If e.g. a sample scatters isotropically over the whole 4π solid angle, and all interesting directions are known to be contained within a certain solid angle interval $\Delta\Omega$, only these solid angles are used for the Monte Carlo choice of scattering direction. According to Eq. (2.9), the weight factor will then have to be changed by the amount $\pi_j = |\Delta\Omega|/(4\pi)$. One thus ensures that the mean simulated intensity is unchanged during a "correct" direction focusing, while a too narrow focusing will result in a lower (i.e. wrong) intensity, since we cut x-rays rays that should have reached the final detector.

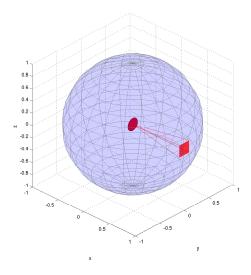


Figure 2.1.: Illustration of the effect of direction focusing in McXtrace. Weights of x-rays emitted into a certain solid angle are scaled down by the full unit sphere area.

2.4. Stratified sampling

One particular efficiency improvement technique is the so-called *stratified sampling*. It consists in partitioning the event distributions in representative sub-spaces, which are then all sampled individualy. The advantage is that we are then sure that each sub-space is well represented in the final integrals. This means that instead of shooting N events, we define D partitions and shoot r = N/D events in each partition. We may define partitions so that they represent 'interesting' distributions, e.g. from events scattered on a monochromator or a sample. The sum of partitions should equal the total space integrated by the Monte Carlo method, and each partition must be sampled randomly.

In the case of McXtrace, the stratified sampling is used when repeating events, i.e. when using the SPLIT keyword in the TRACE section on beamline descriptions. We emphasize here that the number of repetitions r should not exceed the dimensionality of the Monte Carlo integration space (which is d = 10 for x-ray events) and the dimensional strategies of the context of the co

Records	Accurarcy
10^{3}	10 %
10^{4}	2.5~%
10^{5}	1 %
10^{6}	0.25~%
10^{7}	0.05~%

Table 2.1.: Accuracy estimate as a function of the number of statistical events used to estimate an integral with McXtrace.

sionality of the partition spaces, i.e. the number of random generators following the stratified sampling location in the beamline.

2.5. Accuracy of Monte Carlo simulations

When running a Monte Carlo, the meaningful quantities are obtained by integrating random events into a single value (e.g. flux), or onto an histogram grid. The theory [Jam80] shows that the accuracy of these estimates is a function of the space dimension d and the number of events N. For large numbers N, the central limit theorem provides an estimate of the relative error as $1/\sqrt{N}$. However, the exact expression depends on the random distributions.

McXtrace uses a space with d = 12 parameters to describe x-rays (position, wavevector, weight, polarisation, phase, time). We show in Table 2.1 a rough estimate of the accuracy on integrals as a function of the number of records reaching the integration point. This stands both for integrated flux, as well as for histogram bins - for which the number of events per bin should be used for N.

3. Source components

McXtrace contains a number of different source components, and any simulation will usually contain exactly one of these sources. The main function of a source is to determine a set of initial parameters $(\mathbf{r}, \mathbf{k}, t)$ for each photon ray. This is done by Monte Carlo choices from suitable distributions. For example, in most present sources the initial position is found from a uniform distribution over the source surface. The initial photon wavenumber is selected within an interval of either the corresponding energy or the corresponding wavelength.

For pulsed sources, the choice of the emission time, t, is being made on basis of detailed analytical expressions. For other sources, t is set to zero. In the case one would like to use a steady state source with time-of-flight settings, the emission time of each photon ray should be determined using a Monte Carlo choice. This may be achieved by the EXTEND keyword in the instrument description source as in the example below:

```
TRACE

COMPONENT MySource=Source_pt(...) AT (...)

EXTEND

%{
   t = 1e-3*randpm1(); /* set time to +/- 1 ms */

%}
```

Also take a look at the **Chopper_simple** component.

3.0.1. Photon flux and Brilliance

The flux of the sources deserves special attention. The total intensity is defined as the sum of weights of all emitted x-rays during one simulation (the unit of total photon weight is thus photons per second). The flux, ψ , at an instrument is defined as intensity per area perpendicular to the beam direction.

The source Brilliance, Φ , is defined in different units (See e.g. [ANM11]): the number of photon rays emitted per second from a 1 m²m area on the source surface, with direction within a 1 m²rad angle window, and with wavelength within a 1% interval. The total intensity of real photons emitted towards a given diaphragm (units: ph./s) is therefore (for constant Φ):

$$I_{\text{total}} = \Phi A \Delta \Omega \Delta \lambda, \tag{3.1}$$

where A is the source area, $\Delta\Omega$ is the solid angle of the diaphragm as seen from the source surface, and $\Delta\lambda$ is the width of the wavelength interval in which photons are emitted (assuming a uniform wavelength spectrum).

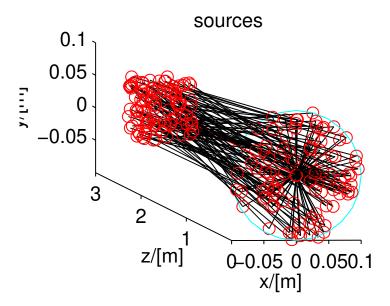


Figure 3.1.: A circular source component (at z=0) emitting photon rays randomly, either from a model, or from a data file.

The simulations are performed so that detector intensities are independent of the number of photon histories simulated (although more photon histories will give better statistics). If N_{sim} denotes the number of x-ray histories to simulate, the initial photon weight p_0 must be set to

$$p_0 = \frac{N_{\text{total}}}{N_{\text{sim}}} = \frac{\Phi(\lambda)}{N_{\text{sim}}} A\Omega \Delta \lambda, \tag{3.2}$$

where the source brilliance is now given a λ -dependence.

As a start, we recommend new McXtrace users to use the **Source_flat** component. For a slightly more realistic sources are supply **Source_flat** with a spectrum file (for instance generated by SPECTRA [TK01]) or **Source_gaussian**.

Optimizers can dramatically improve the statistics, but may occasionally give wrong results, due to misleaded optimization. You should always check such simulations with (shorter) non-optimized ones.

Other ways to speed-up simulations are to read events from a file. See section 3.6 for details.

3.1. Source_pt: A mathematical point emitting photons with a spectrum either uniform, gaussian or generated from a datafile

Input Parameters for component Source_pt from sources

```
<Parameter = value>, [Unit], Description
   <spectrum_file=NULL>: [ ] File from which to read the spectral intensity
  <focus_xw=0>: [m] Width of target
  <focus_yh=0>: [m] Height of target
  <focus_x0=0>: [m] x-cocordinate of target centre.
  <focus_y0=0>: [m] y-coordinate of target centre.
  <flux=0>: [ph/s] Total flux radiated from the source.
   <dist=1>: [m] Distance from source plane to sampling window.
   <E0=0>: [keV] Mean energy of xrays.
  <dE=0>: [keV] Energy half spread of x-rays.
  < lambda0 = 0 >: [AA] Mean wavelength of x-rays.
  <dlambda=0>: [AA] Wavelength half spread of x-rays (flat or gaussian sigma)
13
  <phase=0>: [ ] Set phase to something given.
   <randomphase=1>: [ ] If nonzero, the phase of the emotted photon is random,
       i.e. source is fully incoherent. otherwise the value of phase is used.
  <gauss=0>: [1] Gaussian (1) or Flat (0) energy/wavelength distribution
15
```

The simplest source model, where a mathematical point source at (0,0,0) emits photons. The wavevector of the emitted photons is picked randomly in a defining aperture $focus_xw$ by $focus_yh$ m at (0,0,dist). Please note that this aperture is merely a virtual aperture used to reduce the sampling space. This has a few implications: Other components may be placed without reference to the aperture, but if the aperture does not fill the full acceptance window of subsequent components your simulations will be biased. The aperture is simply there to provide efficient sampling.

If a spectrum_file is not supplied, the xray is given a weight which is the total wavelength-integrated intensity downscaled by the solid angle subtended by the defining aperture. The Energy/wavelength spectrum of the emitted photons is centered around $E\theta$ or $\lambda 0$ with a width of dE or $d\lambda$ respectively. $E\theta$ takes precedence over $\lambda 0$. Thus if $E0 \neq 0$ the the combination $E\theta$, dE is used, $\lambda 0$, $d\lambda$ otherwise. If gauss (the default) is set to be non-zero the spectrum is Gaussian with mean $E\theta$ ($\lambda 0$) and standard deviation dE ($d\lambda$), otherwise the spectrum is uniform in $E\theta \pm dE$ ($\lambda 0 \pm d\lambda$).

If a *spectrum_file is* supplied, a slightly different strategy is adopted. In this case the wavelength/energy range implied by the datafile is sampled unformly and each ray is assigned a weight corresponding to the intensity indicated by linear interpolation between datapoints at that wavelength. This implies an oversampling of weak parts of the intensity spectrum.

Currently only completely coherent or fully incoherent beams are supported. If randomphase is specified emitted photons will be assigned a random phase, otherwise it is set to the value of phase.

3.2. Source_flat: A flat surface emitting photons with a spectrum either uniform, gaussian or generated from a datafile

Input Parameters for component Source_flat from sources

```
1 | < Parameter = value >, [Unit], Description
2 | <spectrum_file=NULL>: [ ] File from which to read the spectral intensity
3 < radius=0>: [m] Radius of circle in (x,y,0) plane where x-rays are
      generated.
  <yheight=0>: [m] Height of rectangle in (x,y,0) plane where x-rays are
      generated.
  <xwidth=0>: [m] Width of rectangle in (x,y,0) plane where x-rays are
5
      generated. Overrides xmin and xmax.
6
  <xmin=0>: [m] Lower bound of x-interval where photons are generated.
  7
9 | < focus_xw = .045 > : [m] Width of target
10 < focus_yh = .12>: [m] Height of target
11 < focus_r=0>: [superseeds focus_xh, focus_yh] [m] Radius of target
12 <E0=0>: [keV] Mean energy of xrays.
13 <dE=0>: [keV] Energy half spread of x-rays (flat or gaussian sigma).
14 < lambda0=0>: [AA] Mean wavelength of x-rays.
15 < dlambda = 0>: [AA] Wavelength half spread of x-rays.
16 < flux = 0>: [pht/s] total flux radiated from the source
17 | <gauss=0>: [1] Gaussian (1) or Flat (0) energy/wavelength distribution
18 < randomphase = 1>: [ ] If nonzero, the phase of the emotted photon is random,
       i.e. source is fully incoherent. otherwise the value of phase is used.
19 | <phase=0>: [ ] Set phase to something given.
```

A simple source model, with a flat surface emitting photons. The surface in the xy-plane is specified as a rectangle with dimensions $xwidth \times yheight$ m², or as a circle with radius m. The initial x-ray position is chosen randomly in the source surface — its wavevector is chosen randomly (exactly as in the case of **Source_pt**) (section 3.1) in the defining aperture with height $focus_yh$ and width $focus_xw$ placed at (0,0,dist).

Just as for **Source_pt** the aperture is for efficiency purposes and, if misused, may cause biasing.

With the exception of source size related parameters, all other parameters are identical to **Source_pt** Note that this also applies to coherence and photon phase. If random-phase=0 then a photon emitted from $(x_0, y_0, 0)$ will in phase with a photon emitted from $(x_1, y_1, 0)$. I.e. full transversal coherence.

3.3. Source_div: A continuous source with specified divergence

Input Parameters for component Source_div from sources

```
1 < Parameter = value >, [Unit], Description
```

```
2 | < spectrum_file=NULL>: [ ] File from which to read the spectral intensity
      profile
3 | < xwidth = 0>: [m] Width of source.
4 < yheight=0>: [m] Height of source.
5 | cfocus_aw=0>: [rd] Standard deviation (Gaussian) or maximal (uniform) horz.
       width divergence.
  <focus_ah=0>: [rd] Standard deviation (Gaussian) or maximal (uniform) vert.
       height divergence.
   <focus_ar=0>: [rd] Standard deviation (Gaussian) or maximal (uniform)
      radial divergence.
   <radius=0>: [m] Radius of circular source
  \langle E0=0 \rangle: [keV] Mean energy of X-rays.
  <dE=0>: [keV] Energy half spread of X-rays. If gauss==0 dE is the half-
      spread, i.e. E \setminus in[E0-dE,E0+dE], if gauss!=0 it's interpreted as the
      standard dev.
11 < lambda0=0>: [AA] Mean wavelength of X-rays (only relevant for E0=0).
12 <dlambda=0>: [AA] Wavelength half spread of X-rays.
13 <flux=0>: [1/(s * mm**2 *mrad**2 * energy unit)] flux per energy unit, Angs
       or keV.
  <gauss=0>: [1] Criterion: 0: uniform, 1: Gaussian distribution of energy/
14
      wavelength.
   <gauss_a=0>: [1] Criterion: 0: uniform, 1: Gaussian divergence distribution
16
  <randomphase=1>: [1] If !=0 the photon phase is chosen randomly.
  <phase=0>: [1] Value of the photon phase (if randomphase==0).
```

Source_div is a rectangular source, $w \times h$, which emits a beam of a specified divergence around the direction of the z-axis.

Just as for **Source_pt**, if a *spectrum_file* is not supplied, the xray is given a weight which is the total wavelength-integrated intensity downscaled by the solid angle subtended by the defining aperture and the energy/wavlength spectrum is centered around E0 or $\lambda0$ with width dE or $d\lambda$ respectively. The profile is Gaussian if $gauss \neq 0$, uniform if gauss = 0, If a $spectrum_file$ is supplied, a slightly different strategy is adopted. In this case the wavelength/energy range implied by the datafile is sampled unformly and each ray is assigned a weight corresponding to the intensity indicated by linear interpolation between datapoints at that wavelength. This implies an oversampling of weak parts of the intensity spectrum.

The beam intensity is uniform over the whole of the source and the source divergences are $focus_ah$ and $focus_aw$ in degrees. Unless $gauss_a \neq = 0$ the individual photons are sampled from a uniform divergence distribution, otherwise a Gaussian is used.

3.4. Source_gaussian: the model has a gaussian distribution of intensity

Input Parameters for component Source_gaussian from sources

```
1 < Parameter = value >, [Unit], Description
```

```
2 | < spectrum_file=NULL>: [ ] File from which to read the spectral intensity
      profile
3 < sig_x=1>: [m] Horizontal standard deviation of source (rms source size).
4 < sig_y=0>: [m] Vertical standard deviation of source (rms source size).
5 < sigPr_x=0>: [rad] Angular horizontal divergence
6 < sigPr_y=0>: [rad] Angular vertical divergence
  <flux=1>: [ ] Scaling factor to set the total emitted unrestricted flux.
8 | <br/> | spectrum_file is Brilliance - apply corrections
       to get to raw flux.
9 < dist=1>: [m] Distance from source plane to sampling window.
10 | <gauss=0>: [ ] Gaussian (1) or uniform (0) spectrum profile.
11 < focus_xw=0>: [m] Width of sampling window dist m downstream from source to
       allow focused sampling.
12 < focus_yh=0>: [m] Height of sampling window dist m downstream from source
      to allow focused sampling.
13 <E0=0>: [overrides spectrum_file] [keV] Centre of emitted energy spectrum
14 <dE=0>: [kev] Half-width (or std. dev.) of emitted energy spectrum.
15 < lambda0=0>: [AA] Centre of emitted wavelength spectrum.
16 < dlambda = -1>: [AA] Half-width (or std. dev.) of emitted wavelength spectrum
17 < phase=0>: [ ] The initial phase of the photons.
18 < randomphase = 1>: [ ] If nonzero phase is random (incoherent radiation),
      otherwise it is set to the value of phase
```

A simplified version of a completely incoherent source of horizontal and vertical sizes $sig_{-}x$ and $sig_{-}y$ respectively with angular divergence $sigPr_{-}x$ and $sigPr_{-}y$. Can be used to model an undulator source emitting a photon beam that has gaussian distribution. This naturally generates a larger Gaussian profile at a distance. A sampling window at (0,0,dist) may be specified by the parameters $focus_{-}xw$ and $focus_{-}yh$, which restricts the sampling phases space of the emitted photons (See section 3.1 for details).

The energy spectrum emitted by **Source_gaussian** source is completely analogous to **Source_pt**, as is its photon phase functionality.

3.5. Source_lab: X-ray tube laboratory source

Input Parameters for component Source_lab from sources

Source_lab is a model of a laboratory X-ray tube. An electron ray hits a target of specified material. Currently, only single material targets are allowed. To model multiple material targets one could construct a model with two or more sources simultaneously. This has consequences for intensity of the source which should be downscaled accordingly.

An electron beam of rectangular transverse crossection (width,height) and energy E0 impinges on the target of material. Wrt. the electron beam, the target is considereded infinitely thick. The beam is considered to have uniform intensity. Thus, the spatial distribution of x-ray generation will be exponential in the depth of the material.

Further, an exit aperture is defined with dimensions (focus_xw,focus_yh). The centre of the aperture is situated at a distance dist m from where the electron beam hits the target slab at an elevation of take_off (see Figure 3.5).

The **Source_lab** coordinate system has its origin in the center of the electron beam at the surface of the anode material and is oriented such that the z-axis points at the center of the exit window, and the x-axis is parallel to the *width* of the electron beam.

Note that the exit aperture is merely an opening. If the material absorption of a window, e.g. Be, is to be taken into account a **Filter** (section 4.5) should be inserted after the exit aperture.

For each photon ray to be generated, a Monte Carlo choice is made to generate either a Bremsstrahlung photon or one from one of the x-ray emission lines of the material. (1 - frac) of the rays are generated from characteristic emission, and (frac) from Bremsstrahlung. In most cases Bremsstrahlung is unwanted background, which is why the default is 0.1. Note that this *only* governs how much of the available statistics is diverted into simulating Bremsstrahlung. It does not have an impact on what intensity is detected in subsequent monitors — only on the errorbars of the detected numbers.

The spectral characteristics of the generated Bremsstrahlung is governed by the model suggested by Kramer [Kra23]. Although disputed in several subsequent papers, the model is simple, and sufficiently accurate for many background estimation purposes.

Characteristic emission on the other hand is sampled from a set of Lorentzian functions with central wavelengths found in the work by [Bea67] with spectral widths taken from [KO79].

An example of beam spectral characteristics emitted from a Cu-anode target detected 1 mm from an exit aperture of 1×1 cm² 10 cm downstream from the target at a $take_off$ angle of 6° is seen in figure 3.5.

Source_lab includes a set of common anode materials: {Cu, Ga, Mo, Ag, W}. More materials can be added by the following procedure:

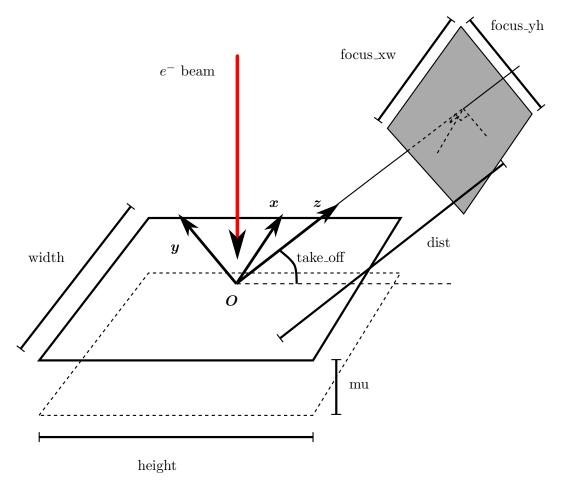


Figure 3.2.: Geometry of the **Source_lab** component. An electron beam impinges at a right angle on an anode material, where X-rays are generated. The Origin is defined to be a the centre of the electron beam on the anode surface, and the coordinate system is oriented such that the **z**-axis point towards the exit aperture.

Figure 3.3.: Intensity vs. wavelenghth for a Cu-anode laboratory source.

- 1. Find the atomic number, Z, for the material you want to add. So far only single materials are supported.
- 2. Look up (and note down) the central energies of (up to) 6 characteristic lines for the material in e.g. [Bea67]. Also note down the number lines you have recorded.
- 3. Look up the natural spectral widths of those lines in [KO79].
- 4. Find the relative intensity of the set of lines. For instance from the x-ray data booklet.
- 5. Note down the ionization energy $E_I(Z)$, and flourescence yield, Y(Z), of the anode material.

3.6. Other sources components: virtual sources (event files)

4. Beam optical components: Arms, slits, filters etc.

This chapter contains a number of optical components used to modify the x-ray beam in various ways, as well as the "generic" component **Arm**.

4.1. Arm: The generic component

Input Parameters for component Arm from optics

```
Parameter = value>, [Unit], Description
```

The component **Arm** is empty; is resembles an optical bench and has no effect on the x-ray. The purpose of this component is only to provide a standard means of defining a local coordinate system within the instrument definition. Other components may then be positioned relative to the **Arm** component using the McXtrace meta-language. The use of **Arm** components in beamline definitions is not required but is recommended for clarity. **Arm** has no input parameters.

The first Arm instance in an instrument definition may be changed into a **Progress_bar** (sec. 9.1) component in order to display simulation progress on the fly, and possibly save intermediate results.

4.2. Slit: A beam defining diaphragm

Input Parameters for component Slit from optics

The component **Slit** is a very simple construction. It sets up an opening at z = 0, and propagates the photon rays onto this plane (by the kernel call PROP_Z0). Photons within the slit opening are unaffected, while all other photons are discarded by the kernel call ABSORB.

By using **Slit**, some photons contributing to the background in a real experiment will be neglected. These are, for instance, the ones that scatter off the inner side of the slit, penetrate the slit material, or clear the outer edges of the slit.

The opening of the slit is determined by specifying either a radius, a width (xwidth) and a height (yheight), or absolute limits in x and y (xmin, xmax, ymin, ymax), in order of precedence. If radius is set, the opening is considered circular.

The slit component can also be used to discard insignificant (*i.e.* very low weight) rays, that in some simulations may be very abundant and therefore time consuming. If the optional parameter cut is set, all x-rays with p < cut are ABSORB'ed. This use is recommended in connection with **Virtual_output**.

Slit may be used to model slit diffraction, thourgh judicious use of the resmapling parameters: $focus_xw, focus_yh, focus_x\theta, focus_y\theta$. The similarity in parameter naming for sources (chapter 3) is intentional as computationally the processes are the same. If present the focus parameters cause the ray to be regarded as a Huygens wave and a new direction going towards the resampling window is computed. The resulting difference in phase gives rise to maxima and minima as expected. This feature should be used with care: the further away from the optical axis, the more statistics and finer binning is needed to get meaningful results. Please refer to [BK+13] for more deatils.

4.3. Slit_N: multiple slits

Input Parameters for component Slit_N from optics

Documentation pending.

4.4. Beamstop: A photon absorbing area

Input Parameters for component Beamstop from optics

```
1 < Parameter = value >, [Unit], Description
```

The component **Beamstop** can be seen as the reverse of the **Slit** component. It sets up an area at the z=0 plane. Photons that hit the plane within this area are ABSORB'ed, while all others are unaffected.

By using this component, some photons contributing to the background in a real experiment will be neglected. These are the ones that scatter off the side of the (real) beamstop, or penetrate the absorbing material. Further, the holder of the beamstop is not simulated.

Beamstop can be either circular or rectangular. The input parameters of **Beamstop** are either height and width (xwidth, yheight) or the four coordinates, (xmin, xmax, ymin, ymax) defining the opening of a rectangle, or the radius of a circle, depending on which parameters are specified.

If the "direct beam" (e.g. after a monochromator or sample) should not be simulated, it is possible to emulate an ideal beamstop so that only the scattered beam is left; without the use of **Beamstop**: This method is useful for instance in the case where only photons scattered from a sample are of interest. The example below removes the direct beam and any background signal from other parts of the beamline

```
COMPONENT MySample=PowderN(...) AT (...)
EXTEND
%{
  if (!SCATTERED) ABSORB;
%}
```

4.5. Filter: A general absorption filter model

Input Parameters for component Filter from optics

This component is a filter in the shape of a rectangular block or a general shape defined by a set of polygons. Given an input file containing material parameters. Neccessary parameters are nominal density and a parameterization of the linear attenuation coefficient, μ as a function of wavelength (or energy).

The model is very simple: Firstly the X-ray is traced to find intersection points between ray and filter (0 or 2). If no intersection is found the x-ray is left untouched and nothing further happens. Assuming the ray intersects the filter: Secondly, the path length dl within the filter is computed. Thirdly a $\mu = f(\lambda, \text{material})$ is computed by interpolating in a datafile, and the x-ray weight is adjusted according to $p = p \exp(-dl * \mu)$. The x-ray is left at the point where it exits the filter block (the 2nd intersection).

Example data files corresponding to all elements up to Z=92 are distributed with McXtrace in the MCXTRACE/data directory as *.txt files. These tables have been extracted from the NIST FFAST [Nis] x-ray database. To generate other datafiles from the same source a simple shell script: MCXTRACE/data/get_xray_db_data is also distributed with McXtrace Running this script will connect to the NIST webiste and download a

.html file. This output must now be modified such that html-tags are removed and all header lines begin with #

4.5.1. Example

This is an example of how to download and generate datafiles for the Filter.comp and others

The distributed tables have been extracted from the NIST x-ray database. To ease generation of more dtafiles from the same source a simple shell script:

```
MCXTRACE/data/get_xray_db_data
```

is also distributed with McXtrace

Running this script will connect to the NIST webiste and download a .html file. This output must now be modified wuch that html-tags are removed and all header lines begin with #.

```
/usr/local/lib/mcxtrace/data/get_xray_db_data 3 output.dat
```

where the second parameter (3) is the atom number of the material, for which we want to generate a datafile. Now open the generated datafile (output.dat) with your favourite text editor and make sure the file ends up looking like this

```
#Li (Z 3)
#Atomic weight: A[r] 6.941000
#Nominal density: rho 5.3300E-01
# rho[a](barns/atom) = [mu/rho](cm^2 g^-1) x 1.15258E+01
              E(eV) \ [mu/rho] \ (cm^2 g^-1) = f[2] \ (e \ atom^-1) \ x \ 6.06257E+062 \ edges. Edge \ energies \ (keV): 
                                 5.47500E-02 L I 5.34000E-03
#Relativistic correction estimate f[rel] (H82,3/5CL) = -9.8613E-04,
                  6.0000E-04 e atom^-1
             Nuclear Thomson correction f[NT] = -7.1131E-04 e atom^-1
#Form Factors, Attenuation and Scattering Cross-sections
\#Z=3, E = 0.001 - 433 keV
           E
                                      f[1]
                                                                                                                  [mu/rho]
                                                                                                                                                     [sigma/rho] [mu/rho] [mu/rho] [K] lambda
                                                                           f[2]
                                                                                                          Photoelectric Coh+inc
1 cm^2 g^-1 cm^2 g^-1
                                                                                                                                                                                           Total
cm^2 g^-1
                                           e atom^-1
                                                                                e atom^-1
       keV
                                                                                                                                                                                                                              cm^2 g^-1 nm
# KeV e atom -1 e atom -1 cm zg -1 cm z
5.334660E-03 8.03599E-01 0.0000E+00 0.0000E+00 2.5813E-07 2.5813E-07 0.000E+00 2.324E+02
5.366700E-03 8.56971E-01 1.0769E-01 1.2165E+05 2.6435E-07 1.2165E+05 0.000E+00 2.310E+02
3.788588E+02 3.00000E+00 3.9121E-08 6.2602E-07 8.4389E-02 8.4390E-02 6.123E-07 3.273E-03
4.050001E+02 3.00000E+00 3.3438E-08 5.0054E-07 8.2127E-02 8.2128E-02 4.895E-07 3.061E-03 4.329451E+02 3.00000E+00 2.8581E-08 4.0022E-07 7.9892E-02 7.9892E-02 3.913E-07 2.864E-03
                                                                                                                                                                                      7.9892E-02 3.913E-07 2.864E-03
```

Please make sure you don't forget to remove the html-tags in the bottom of the file as well. In the future we will set up a more streamlined way of doing this.

4.6. Chopper_simple: An ideal chopper

Input Parameters for component Chopper_simple from optics

```
Parameter = value>, [Unit], Description

2 <t0=0>: [s] Initial delay of the opening time

3 <T=1>: [s] Period of the chopper

4 <tau=0.1>: [s] Opening time of the chopper

5 <xwidth=0.1>: [m] height of the chopper opening

6 <yheight=0.1>: [m] width of the chopper opening

7 <isfirst=0>: [] is the chopper the first chopper on a continous source.

8 <t_rise=0>: [s] Rise time of the chopper pulse.

9 <tjit=0>: [] Timing jitter in terms of the opening time tau. For each ray the opening window will be shifted by a random amount within t=[-tjit *.5*tau, tjit*.5*tau]
```

Chopper_simple is an idelized version of a chopper with a rectangular chopper opening which may open instantly and has no side-scattering etc.

It models the chopper with a blocking infinitely thin aperture which becomes transparent in the time interval $t \in [t\theta, t\theta + \tau]$. This is an idelized version of a chopper where the chopper opening is rectangular which may open instantly (if $t_rise = 0$, the default). For nonzero rise time the aperture simply becomes gradually less opaque for $t \in [t\theta - t_rise, t\theta]$.

For correct normalization of intesity a chopper period, T, must also be set.

is_first is useful when using **Chopper_simple** with continous sources who inherently have no time-dependence. Thus the emission time of the photon ray is arbitrary, and the chopper defines the temporal signature of the beam, i.e. it simply sets the time-parameter of the photon ray randomly in the opening window of the chopper. Naturally this should only be used for the *first* chopper element in a simulation.

5. Refractive optical components: lenses

An X-ray refractive lens, often referred to as a Compund Refractive Lens (CRL), is a fairly new type of device, which has gained popularity in the last few years. An early study showing the feasibility of such devices may be found in [SKS96]. As the refractive index of X-rays is $n \approx 1$ a number of lenses, stacked together is usually necessary to bring the focal length to practical values. McXtrace includes a few lens components, which all have slightly different characterities.

5.1. Lens_simple: Thin lens approximation

Input Parameters for component Lens_simple from optics

This models a thin-lens approximation of a stack of parabolic refractive x-ray lenses

5.2. Lens_parab: Thick parabolic CRL

Input Parameters for component Lens_parab from optics

```
8 | <rough_z=0>: [rad] RMS value of random slope error along z.
9 | <rough_xy=0>: [rad] RMS value of random slope error along x and y.
```

Component model of a stack of compund refractive lenses. Each lens in the stack is modelled by two 2D parabolic surfaces, and rays are traced through all the complete stack taking the displacement of the surfaces into account. This is naturally less efficient than a thin lens approximation.

The logic behind the model is the following: a photon interacts with the surface of the lens at a certain angle Θ , which alters in accordance with Snell's law upon photon's entering the lens's material. The combination of the refractive process inside material (that is characterized by a material datafile) and the interaction with a geometrical surface results in a photon's new trajectory, i.e. in focusing.

The functionality of Lens_parab_Cyl, Lens_parab_rough, and Lens_parab_Cyl_rough will be merged into this component.

5.3. Lens_parab_Cyl: Thick 1D-parabolic CRL

Input Parameters for component Lens_parab_Cyl from optics

This component is based on the same logical approach as the **Lens_parab** with one significant difference - geometrical surface. Parabolic cylinder (parabolic curvature along the vertical axes only, invariant along the horizontal) and thus ofocuses the beam in only in the vertical direction.

This component and its functionality is scheduled to be merged into Lens_parab

5.4. Lens_parab_rough: Thick parabolic CRL including roughness-model

Input Parameters for component Lens_parab_rough from optics

```
1 <Parameter = value >, [Unit], Description
2 <material_datafile="Be.txt">:
3 <r=0.5e-3>: [-] - raduis of curvature (circular approximation at the tip of the profile) [mm];
```

Identical to **Lens_parab** except it has the option of a *roughness* parameter. Roughness is simply modelled by a stochatstic, normally distributed, displacement of the normal vector of the lens surfaces.

This component and its functionality is scheduled to be merged into Lens_parab

5.5. Lens_parab_Cyl_rough: Thick 1D-parabolic CRL including roughness-model

Input Parameters for component Lens_parab_Cyl_rough from optics

Identical to **Lens_parab_Cyl** except it has the option of a *roughness* parameter. Roughness is simply modelled by a stochatstic, normally distributed, displacement of the normal vector of the lens surfaces.

This component and its functionality is scheduled to be merged into Lens_parab

5.6. Lens Kinoform: refractice kinoform lens

Input Parameters for component Lens_Kinoform from optics

```
1 < Parameter = value >, [Unit], Description
2 < yheight=1e-2>:
3 < xwidth=5.268e-4>:
4 < deltaN=0>:
5 < datafile="kinoform.txt">:
```

```
6 | <material_datafile="Si.txt">:
7 | <Re=2.8179402894e-5>:
8 | <Na=6.02214179e23>:
```

Doc. Pend.

5.7. Lens_elliptical:

Input Parameters for component Lens_elliptical from optics

doc. pend.

6. Reflective optical components: mirrors

This section describes advanced reflective X-ray optics components such as mirrors.

6.1. Mirror_curved: Cylindrically curved mirror

Input Parameters for component Mirror_curved from optics

Models a cylindrical mirror, positioned in the XZ-plane curving towards positive X. The input parameter radius defines the radius of curvature and the mirror size is given by length and width where length and width is along Z and Y respectively. coating and $R\theta$ are mutually exclusive. If $R\theta$ is nonzero, it is taken as the reflectivity value, irrespective of wavelength, whereas coating nominates a file from which to read values for f_1 and f_2 . See [Nis] for definitions. For elements $Z \in [1,92]$ datafiles are distributed with the McXtrace system that may be used as: coating="Rh.txt". or coating="Al.txt".

This component is scheduled to be merged with Mirror_parabolic and Mirror_elliptic

6.2. Mirror_parabolic: Mirror with a parabolic curvature profile.

Input Parameters for component Mirror_parabolic from optics

This component is scheduled to be merged with Mirror_elliptic

6.3. Mirror_elliptic: Mirror with a elliptic curvature profile.

Input Parameters for component Mirror_elliptic from optics

```
Parameter = value>, [Unit], Description

(x_a=0>: [m] 1st short half axis (along x). Commonly set to zero, which
    really implies infinite value, so crystal is an elliptic cylinder.

(y_b=1.0>: [m] 2nd short half axis (along y), which is also the presumed
    near-normal direction, reflection near the y-z plane.

(z_c=1.0>: [m] long half axis (along z). Commonly a=0. b=c, which creates
    a circular cylindrical surface.

(length=0.05>: [m] z depth (length) of the crystal.

(width=0.02>: [m] x width of the crystal.

(R0=1>: [1] Reflectivity of mirror.

(coating="">: [str] Datafile containing either mirror material constants or
    reflectivity numbers.
```

This component is scheduled to be merged with Mirror_parabolic

6.4. Multilayer_elliptic: Elliptically curved mirror coated with a multilayer

Input Parameters for component Multilayer_elliptic from optics

```
1 | < Parameter = value >, [Unit], Description
2 < reflectivity_datafile="reflectivity.txt">: [ ] Datafile containing
       reflectivity values as a function of q and E.
3 < theta=1.2>: [deg.] Design angle of incidence.
4 < s1>: [m] Design distance from the source to the multilayer.
5 \left| \langle s2 \rangle \right|: [m] Design focusing distance of the multilayer.
6 < length = 0.5 > : [m] Length of the mirror.
  <width=0.2>: [m] Width of the mirror along x-axis.
8
  < R0 = 1 > :
  <Emin=-1>: [keV] Lower limit of energy interval in datafile. Overrides what
       's written in the datafile header.
  <Emax=-1>: [keV] Upper limit of energy interval in datafile. Overrides what
       's written in the datafile header.
  <Estep=-1>: [keV] Step between energy sample points in datafile. Overrides
      what's written in the datafile header.
12 | <Gamma=0>: [ ] High electron density fraction of bilayer (in kinematical
      appr.).
13 < Lambda=0>: [m] Thickness of bilayer (in kinematical appr.).
  <rho_AB=0>: [ ] Number electron density constrast in bilayer (in
       kinematical appr.).
  <N=0>: [ ] Number of bilayers (in kinematical appr.).
```

The component **Multilayer_elliptic** models a single rectangular reflecting multilayer mirror plate with elliptical curvature. It can be used as a sample component, to *e.g.* assemble a Kirkpatrick-Baez focusing system or in combination with a double-crystal monochromator.

Figure 6.1*Left* shows a side view of a mirror (the blue section of the ellipse) in the McXtrace coordinate system. At the mirror center, the mirror tangent is parallel to the z axis and the mirror normal is parallel to the y axis. The width of the mirror is w and in y-z plane the mirror has the curvature of an ellipse with major axis a and minor axis a.

$$\frac{z^2}{a^2} + \frac{y^2}{b^2} = 1, |x| < \frac{w}{2}. \tag{6.1}$$

The length of the mirror is L. The coordinates of the mirror center $(0, Y_0, Z_0)$ and the ellipse parameters a, b are determined uniquely by the central glancing angle, the source-mirror distance and the mirror-image distance. The position of the mirror is chosen to be at the positive side of the y axis.

The input parameters of this component are: theta [°], the incident angle; s1 [m], the distance from the source to the multilayer; s2 [m], the focusing distance of the multilayer; length [m], the length of the mirrors; width [m], the width of the mirror along the x-axis; R, the reflectivity.

6.4.1. Definition of the reference frames

The direction and position of the incoming photon is defined relative to the coordinate system illustrated in Fig. 6.1Left (in the code referred to as $McXtrace\ coordinate\ system$):

- the y-axis is parallel to the central mirror normal
- the z-axis is parallel to the central mirror tangent
- the origin is at the mirror center

However, all the calculations are conducted in another reference frame which is illustrated in Fig. 6.1 Right(in the following referred to as the Ellipse coordinate system):

- the z-axis is parallel to major axis of ellipse
- the y-axis is parallel to minor axis of ellipse
- the origin is at the center of the ellipse
- the mirror center at $(0, Y_0, Z_0)$, uniquely determined by the glancing angle at the mirror center, the source-mirror distance and mirror-image distance.

6.4.2. Algorithm

- 1. The photon is generated with a starting point S and a direction V_{in} defined in the McXtrace coordinate system.
- 2. All calculations are performed in the *Ellipse* coordinate system, so to proceed the basis is changed to that reference frame.

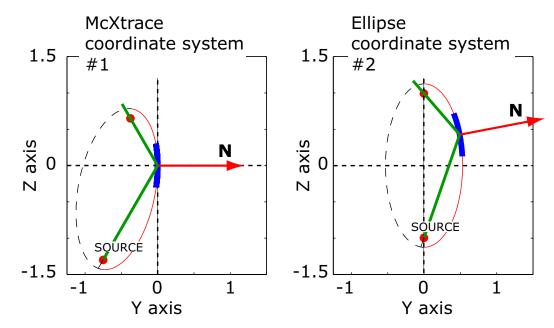


Figure 6.1.: The same image in different coordinate systems.

Left: McXtrace System with the y-axis is parallel to the central mirror normal, the z-axis is parallel to the central mirror tangent and the origin is at the mirror center.

Right: Ellipse System with the z-axis parallel to major axis of ellipse, the y-axis is parallel to minor axis of ellipse and the origin is at the center of the ellipse.

- 3. The 2 intersections of the ray with the ellipse are determined.
- 4. It is checked if any of the intersections are within the area defined by the mirror.
- 5. If one of the solutions is valid, the reflection of that ray is determined.
- 6. The coordinates of the starting point and direction of the reflected ray are calculated using the basis of the *McXtrace* coordinate system.

6.5. Reflection of the ray in the mirror

The tangent and normal to the ellipse $z^2/a^2+y^2/b^2=1$ at the point (Y,Z) are found by implicit differentiation:

$$\frac{2z}{a^2} + \frac{2y}{b^2} \frac{dy}{dz} = 0, (6.2)$$

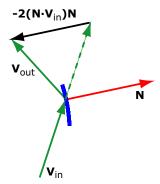


Figure 6.2.: The reflection of the unit vector $\mathbf{V}_{\rm in}$ in the mirror with the normal unit vector \mathbf{N} is $\mathbf{V}_{\rm out} = \mathbf{V}_{\rm in} - 2(\mathbf{N} \cdot \mathbf{V}_{\rm in})\mathbf{N}$

so at the point (Y, Z) the slope of the tangent is $\frac{dy}{dz} = -\frac{Zb^2}{Ya^2}$. The slope of the normal is minus the inverse of the tangent slope, so the coordinates of the mirror normal are

$$N_x = 0$$
 $N_y = \frac{a^2 Y}{b^2 Z}$ $N_z = 1$. (6.3)

With $V_{\rm in}$ and N denoting unit vectors (direction and normal respectively), the direction of the reflected ray is calculated as

$$\boldsymbol{V}_{\text{out}} = \boldsymbol{V}_{\text{in}} - 2(\boldsymbol{N} \cdot \boldsymbol{V}_{\text{in}}) \boldsymbol{N} = \begin{pmatrix} V_{\text{in}x} - 2(\boldsymbol{N} \cdot \boldsymbol{V}_{\text{in}}) N_x \\ V_{\text{in}y} - 2(\boldsymbol{N} \cdot \boldsymbol{V}_{\text{in}}) N_y \\ V_{\text{in}z} - 2(\boldsymbol{N} \cdot \boldsymbol{V}_{\text{in}}) N_z \end{pmatrix}$$
(6.4)

6.5.1. Mirror reflectivity

At present, the Multilayer_elliptic Mirror component uses a reflectivity table reflect, which 1st column is q $[\mathring{A}^{-1}]$ and from the 2nd column on as the reflectivity R in [0-1] as function of tabulated energy [keV]. An example file, calculated for a particular Si/W multilayer, is provided (reflectivity.txt). User provided reflectivity data files can be parsed by the component.

6.6. TwinKB_ML: Side-by-side Kirkpatrick-Baez mirror pair

Input Parameters for component TwinKB_ML from optics

```
 \begin{array}{c} \text{7} < \text{width=0.2>: [m] Width of the mirror along x-axis} \\ \text{8} < \text{R0=0>: [] Constant reflectivity, R0=1 for an ideal situation. If R0=0,} \\ \text{the code reads the reflectivity from the datafile} \\ \end{array}
```

Models a pair of perpendicular, elliptically curved mirrors, known as a Montel-mirror or Side-by-side Kirkpatrick-Baez mirror.

7. Samples

This class of components models the sample of the experiment. This is by far the most challenging part of an xray scattering instrument to model. However, for purpose of simulating instrument performance, details of the samples are rather unimportant, allowing for simple approximations. On the contrary, for full virtual experiments it is of importance to have realistic and detailed sample descriptions. McXtrace contains both simple and detailed samples.

An important component class is elastic Bragg scattering from an ideal powder. The component **PowderN** models a powder scatterer with reflections given in an input file. The component includes absorption, incoherent scattering, direct beam transmission and can assume *concentric* shape, i.e. can be used for modelling sample environments.

Next type is Bragg scattering from single crystals. Two types of single crystal exist in McXtrace at present: The Perfect_crystal and Single_crystal components. Perfect_crystal is in fact most often used as a monchromator crystal. It models a crystal where peak broadening is dominated by the Darwin width. Currently it only handles a single defined reflection. If more than one is wanted this could be accomplished by using two instances in a GROUP and dynamically choose between them with a WHEN-statement. For details on the GROUP and WHEN constructs see the main McXtrace user manual [Knu+14]

Much more general, the component **Single_crystal** is a single crystal sample (with multiple scattering) that allows the input of an arbitrary unit cell and a list of structure factors, read from a LAZY / Crystallographica file. This component also allows anisotropic mosaicity and $\Delta d/d$ lattice space variation.

Isotropic small-angle scattering is simulated in **Saxs_Spheres**, which models scattering from a collection of hard spheres (dilute colloids). Furthermore, a whole series of sample components modelling various SAXS standard sample types are available in the contrib component library section.

7.0.1. Scattering notation

In sample components, we use a notation common for scattering experiments, where the wave vector transfer is denoted the *scattering vector*

$$q \equiv k_{\rm i} - k_{\rm f}. \tag{7.1}$$

In analygo, the *energy transfer* is given by

$$\hbar\omega \equiv E_{\rm i} - E_{\rm f} = \frac{\hbar^2}{2m_{\rm n}} \left(k_{\rm i}^2 - k_{\rm f}^2\right). \tag{7.2}$$

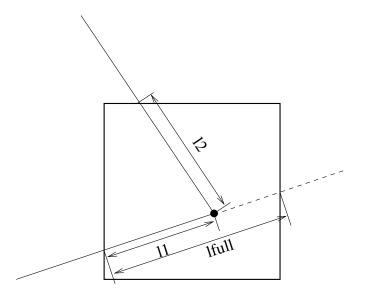


Figure 7.1.: The geometry of a scattering event within a powder sample.

7.0.2. Weight transformation in samples; focusing

Within many samples, the incident beam is attenuated by scattering and absorption, so that the illumination varies considerably throughout the sample. For single crystals, this phenomenon is known as *secondary extinction* [Bac75], but the effect is important for all samples. In analytical treatments, attenuation is difficult to deal with, and is thus often ignored, making a *thin sample approximation*. In Monte Carlo simulations, the beam attenuation is easily taken care of, as will be shown below. In the description, we ignore multiple scattering, which is however implemented in some sample components.

The sample has an absorption cross section per unit cell of σ_c^a and a scattering cross section per unit cell of σ_c^s . The x-ray path length in the sample before the scattering event is denoted by l_1 , and the path length within the sample after the scattering is denoted by l_2 , see figure 7.1. We then define the inverse penetration lengths as $\mu^s = \sigma_c^s/V_c$ and $\mu^a = \sigma_c^a/V_c$, where V_c is the volume of a unit cell. Physically, the attenuation along this path follows

$$f_{\text{att}}(l) = \exp(-l(\mu^s + \mu^a)),$$
 (7.3)

where the normalization $f_{\text{att}}(0) = 1$.

The probability for a given x-ray to be scattered from within the interval $[l_1; l_1 + dl]$ will be

$$P(l_1)dl = \mu^s f_{\text{att}}(l_1)dl, \tag{7.4}$$

while the probability for a x-ray to be scattered from within this interval into the solid angle Ω and not being scattered further or absorbed on the way out of the sample is

$$P(l_1, \Omega)dld\Omega = \mu^s f_{\text{att}}(l_1) f_{\text{att}}(l_2) \gamma(\Omega) d\Omega dl, \tag{7.5}$$

where $\gamma(\Omega)$ is the directional distribution of the scattered x-rays, and l_2 is determined by Monte Carlo chocies of l_1 , Ω , and from the sample geometry, see e.g. figure 7.1.

In our Monte-Carlo simulations, we may choose the scattering parameters by making a Monte-Carlo choice of l_1 and Ω from a distribution different from $P(l_1, \Omega)$. By doing this, we must adjust π_i according to the probability transformation rule (2.9). If we e.g. choose the scattering depth, l_1 , from a flat distribution in $[0; l_{\text{full}}]$, and choose the directional dependence from $g(\Omega)$, we have a Monte Carlo probability

$$f(l_1, \Omega) = g(\Omega)/l_{\text{full}},\tag{7.6}$$

 l_{full} is here the path length through the sample as taken by a non-scattered x-ray (although we here assume that all simulated x-rays are being scattered). According to (2.9), the x-ray weight factor is now adjusted by the amount

$$\pi_i(l_1, \Omega) = \mu^s l_{\text{full}} \exp\left[-(l_1 + l_2)(\mu^a + \mu^s)\right] \frac{\gamma(\Omega)}{q(\Omega)}.$$
 (7.7)

In analogy with the source components, it is possible to define "interesting" directions for the scattering. One will then try to focus the scattered x-rays, choosing a $g(\Omega)$, which peaks around these directions. To do this, one uses (7.7), where the fraction $\gamma(\Omega)/g(\Omega)$ corrects for the focusing. One must choose a proper distribution so that $g(\Omega) > 0$ in every interesting direction. If this is not the case, the Monte Carlo simulation gives incorrect results. All samples have been constructed with a focusing and a non-focusing option.

7.0.3. Future development of sample components

There is still room for much more development of functionality in McXtrace samples.

7.1. Absorption_sample: An absorption phantom

Input Parameters for component Absorption_sample from samples

This component models an absorption phantom with a single inclusion in surrounding material. It is intended use is for tomographic imaging simulations.

7.2. Saxs_spheres: A model of dilute hard spheres in solution for SAXS-use

Input Parameters for component Saxs_spheres from samples

```
1 | < Parameter = value >, [Unit], Description
2 sphere_mtrl=NULL>: Material datafile from which to find absorption. If
      none is given absorption is neglected.
3 <R=100>: [AA] Radius of scattering hard spheres
4 < Phi=1e-3>: [1] Particle volume fraction
5 < Delta_rho = 0.6 >: [fm/AA^3] Excess scattering length density
6 < xwidth=0>: [m] horiz. dimension of sample, as a width
7 < yheight=0>: [m] vert . dimension of sample, as a height for cylinder/box
8 < zdepth=0>: [m] depth of sample
9 < radius=0>: [m] Outer radius of sample in (x,z) plane for cylinder/sphere
10 < t \operatorname{arget}_{-x} = 0 > : [-]
11 < target_y=0>: [m] position of target to focus at
12 < \text{target_z} = 6 > : [-]
13 < target_index = 0>: [1] Relative index of component to focus at, e.g. next is
14 <focus_xw=0>: [m] horiz. dimension of a rectangular area
15 < focus_yh=0>: [m] vert. dimension of a rectangular area
16 <focus_aw=0>: [deg] horiz. angular dimension of a rectangular area
17 < focus_ah=0>: [deg] vert. angular dimension of a rectangular area
18 < focus_r=0>: [m] Detector (disk-shaped) radius
19 <mu_c=0>: [5] Column of the datafile which contains absorption coefficients
```

7.3. PowderN: A general powder sample

Input Parameters for component PowderN from samples

```
1 | < Parameter = value >, [Unit], Description
 2 <format=Undefined>: [ ] Name of the format, or list of column indexes (see Description). N.b. no qoutes!
 3 <mat_format=mat_Undefined >:
 4 < reflections = 0>: Input file for reflections.
 5 Use only incoherent scattering if NULL or "" [string]
 6 < material = 0>:
 7 | <geometry=0>: [str] Name of an Object File Format (OFF) or PLY file for
            complex geometry. The OFF/PLY file may be generated from XYZ
            coordinates using qhull/powercrust.
    <radius=0>: [m] Outer radius of sample in (x,z) plane
    <yheight=0>: [m] Height of sample y direction
10 < xwidth=0>: [m] Horiz. dimension of sample, as a width
11 | <zdepth=0>: [m] Depth of box sample
12 < thickness = 0>: Thickness of hollow sample.
13 Negative value extends the hollow volume outside of the box/cylinder. [m]
14 | <pack=1>: [1] Packing factor
     <Vc=0>: [AA^3] Volume of unit cell=nb atoms per cell/density of atoms.
15
     <Delta_d=0>: [AA] Global relative Delta_d/d spreading when the 'w' column
             is not available. Use 0 if ideal.
17 | <p_inc=0.1>: [1] Fraction of incoherently scattered rays.
     <p_transmit=0.1>: [1] Fraction of transmitted (only attenuated) rays.
18
19 Description of the distribution of the dist
             available. Use 1 if included in F2.
20 <nb-atoms=1>: [1] Number of sub-unit per unit cell, that is ratio of sigma
            for chemical formula to sigma per unit cell.
21 < d-phi=0>: [deg] Angle corresponding to the vertical angular range to focus
              to, e.g. detector height. 0 for no focusing.
22 < d_theta = 360>: [deg] Longitudinal angle to focus to, i.e. lateral opening
            angle. This means, the maximum scattering angle, 2*theta, is equal to
            d_{theta}/2.0
23 < p_interact = 0.8 >: [1] Fraction of events interacting with sample, e.g. 1-
            p_transmit-p_inc.
24 | <concentric=0>: [1] Indicate that this component has a hollow geometry and
            may contain other components. It should then be duplicated after the
            inside part (only for box, cylinder, sphere).
25 < density =0>: [g/cm^3] Density of material. rho=density/weight/1e24*N_A.
     <weight=0>: [g/mol] Atomic/molecular weight of material [g/mol]
27 \mid \text{cbarns=1}: [1] Flag to indicate if \mid F \mid^2 from 'reflections' is in barns or
            fm<sup>2</sup>, (barns=1 for laz, barns=0 for lau type files).
     <focus_flip=0>: [1] Controls the sense of d_phi. If 0 d_phi is measured
            against the xz-plane. If !=0 d_phi is measured against zy-plane.
```

The powder diffraction component **PowderN** models a powder sample with background coming only from incoherent scattering and no multiple scattering. At the users choice, a given percentage of the incoming events may be transmitted (attenuated) to model the direct beam. The component can also assume *concentric* shape, i.e. be used for describing sample environment (cryostat, sample container etc.).

The description of the powder comes from a file in one of the standard output formats LAZY, FULLPROF, or CRYSTALLOGRAPHICA.

7.3.1. Files formats: powder structures

Data files of type lau and laz in the McXtrace distribution data directory are self-documented in their header.

```
PowderN(<geometry parameters>, filename="Al.laz")
```

Other column-based file formats may also be imported e.g. with parameters such as:

```
format=Crystallographica
format=Fullprof
format={1,2,3,4,0,0,0,0}
```

In the latter case, the indices define order of columns parameters multiplicity, lattice spacing, F^2 , Debye-Waller factor and intrinsic line width.

The column signification may as well explicitely be set in the data file header using any of the lines:

Other component parameters may as well be specified in the data file header with lines e.g.:

```
#V_rho
              <value of atom number density [at/Angs^3]>
#Vc
              <value of unit cell volume Vc [Angs^3]>
#sigma_abs
              <value of Absorption cross section [barns]>
#sigma_inc
              <value of Incoherent cross section [barns]>
#Debye_Waller <value of Debye-Waller factor DW>
#Delta_d/d
              <value of Detla_d/d width for all lines>
#density
              <value of material density [g/cm^3]>
              <value of material molar weight [g/mol]>
#weight
#nb_atoms
              <value of number of atoms per unit cell>
```

Further details on file formats are available in the mxdoc page of the component.

7.3.2. Geometry, physical properties, concentricity

The sample has the shape of a solid cylinder, radius r and height h or a box-shaped sample of size $xwidth \times yheight \times zdepth$. At the users choice, an inner 'hollow' can be specified using the parameter thickness.

PowderN can assume a *concentric* shape, i.e. can contain other components inside the inner hollow. To allow this, two almost identical copies of the PowderN components must be set up *around* the internal component(s), for example:

```
COMPONENT Cryo = PowderN(reflections="Al.laz", radius = 0.01, thickness = 0.001, concentric = 1)
```

AT (0,0,0) RELATIVE Somewhere

COMPONENT Sample = some_other_component(with geometry FULLY enclosed in the hollow)
AT (0,0,0) RELATIVE Somewhere

```
COMPONENT Cryo2 = COPY(Cryo)(concentric = 0)
AT (0,0,0) RELATIVE Somewhere
```

As outlined, the first instance of PowderN *must* have concentric = 1 and the second instance *must* have concentric = 0. Furthermore, the component(s) inside the hollow *must* have a geometry which can be fully contained inside the hollow.

In addition to the coherent scattering specified in the **reflections** file, absorptionand incoherent cross sections can be given using the input parameters σ_c^a and σ_i^s .

The Bragg scattering from the powder, σ_c^s is calculated from the input file, with the parameters Q, $|F(Q)|^2$, and j for the scattering vector, structure factor, and multiplicity, respectively. The volume of the unit cell is denoted Vc, while the sample packing factor is f_{pack} .

Focusing is performed by only scattering into one angular interval, $d\phi$ of the Debye-Scherrer circle. The center of this interval is located at the point where the Debye-Scherrer circle intersects the half-plane defined by the initial velocity, v_i , and a user-specified vector, f.

7.3.3. Powder scattering

An ideal powder sample consists of many small crystallites, although each crystallite is sufficiently large not to cause measurable size broadening. The orientation of the crystallites is evenly distributed, and there is thus always a large number of crystallites oriented to fulfill the Bragg condition

$$n\lambda = 2d\sin\theta,\tag{7.8}$$

where n is the order of the scattering (an integer), λ is the x-ray wavelength, d is the lattice spacing of the sample, and 2θ is the scattering angle, see figure 7.2. As all crystal orientations are realised in a powder sample, the x-rays are scattered within a Debye-Scherrer cone of opening angle 4θ [Bac75].

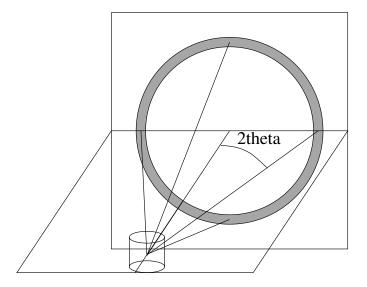


Figure 7.2.: The scattering geometry of a powder sample showing part of the Debye-Scherrer cone (solid lines) and the Debye-Scherrer circle (grey).

Equation (7.8) may be cast into the form

$$|\mathbf{Q}| = 2|\mathbf{k}|\sin\theta,\tag{7.9}$$

where \mathbf{Q} is a vector of the reciprocal lattice, and \mathbf{k} is the wave vector of the x-ray. It is seen that only reciprocal vectors fulfilling $|\mathbf{Q}| < 2|\mathbf{k}|$ contribute to the scattering. For a complete treatment of the powder sample, one needs to take into account all these \mathbf{Q} -values, since each of them contribute to the attenuation.

The strength of the Bragg reflections is given by their structure factors

$$\left| \sum_{j} b_{j} \exp(\mathbf{R}_{j} \cdot \mathbf{Q}) \right|^{2}, \tag{7.10}$$

where the sum runs over all atoms in one unit cell. This structure factor is non-zero only when Q equals a reciprocal lattice vector.

The textbook expression for the scattering cross section corresponding to one Debye-Scherrer cone reads [Squ78, ch.3.6], with $V = NV_0$ being the total sample volume:

$$\sigma_{\text{cone}} = \frac{V}{V_0^2} \frac{\lambda^3}{4\sin\theta} \sum_{Q} |F(Q)|^2.$$
 (7.11)

For our purpose, this expression should be changed slightly. Firstly, the sum over structure factors for a particular Q is replaced by the sum over essentially different reflections multiplied by their multiplicity, j. Then, a finite packing factor, f, is defined for the

powder, and finally, the Debye-Waller factor is multiplied on the elastic cross section to take lattice vibrations into account (no inelastic background is simulated, however). We then reach

$$\sigma_{\text{cone,Q}} = j_Q f \exp(-2W) \frac{V}{V_0^2} \frac{\lambda^3}{4 \sin \theta} |F(Q)|^2$$
(7.12)

$$= f \exp(-2W) \frac{N}{V_0} \frac{4\pi^3}{k^2} \frac{j_Q |F(Q)|^2}{Q}$$
 (7.13)

in the thin sample approximation. For samples of finite thickness, the beam is being attenuated by the attenuation coefficient

$$\mu_{\mathcal{Q}} = \sigma_{\text{cone},\mathcal{Q}}/V. \tag{7.14}$$

For calibration it may be useful to consider the total intensity scattered into a detector of effective height h, covering only one reflection [Squ78, ch.3.6]. A cut though the Debye-Scherrer cone perpendicular to its axis is a circle. At the distance r from the sample, the radius of this circle is $r \sin(2\theta)$. Thus, the detector (in a small angle approximation) counts a fraction $h/(2\pi r \sin(2\theta))$ of the scattered x-rays, giving a resulting count intensity:

$$I = \Psi \sigma_{\text{cone,Q}} \frac{h}{2\pi r \sin(2\theta)},\tag{7.15}$$

powder line.

where Ψ is the flux at the sample position.

For clarity we repeat the meaning and unit of the symbols:

Ψ	${ m s}^{-1}{ m m}^{-2}$	Incoming intensity of x-rays
I	s^{-1}	Detected intensity of x-rays
h	m	Height of detector
r	m	Distance from sample to detector
f	1	Packing factor of the powder
j	1	Multiplicity of the reflection
V_0	m^3	Volume of unit cell
$ F(Q) ^2$	m^2	Structure factor
$\exp(-2W)$	1	Debye-Waller factor
$\mu_{ m Q}$	m^{-1}	Linear attenuation factor due to scattering from one

A powder sample will in general have several allowed reflections Q_j , which will all contribute to the attenuation. These reflections will have different values of $|F(Q_j)|^2$ (and hence of Q_j), j_j , $\exp(-2W_j)$, and θ_j . The total attenuation through the sample due to scattering is given by $\mu^s = \mu^s_{\rm inc} + \sum_j \mu^s_j$, where $\mu^s_{\rm inc}$ represents the incoherent scattering.

7.3.4. Algorithm

The algorithm of **PowderN** can be summarized as

• Check if the x-ray intersects the sample (otherwise ignore the following).

- \bullet Calculate the attenuation coefficients for scattering and absorption.
- Perform Monte Carlo choices to determine the scattering position, scattering type (coherent/incoherent), and the outgoing direction.
- Perform the necessary weight factor transformation.

7.4. Perfect_crystal: A Darwin-width domniated single crystal model

Input Parameters for component Perfect_crystal from samples

Pedning Further Documetation

7.5. Single_crystal: The single crystal component

Input Parameters for component Single_crystal from samples

```
1 | < Parameter = value >, [Unit], Description
 2 \mid < mosaic\_AB = \{0>: [arc\_minutes, arc\_minutes, 1, 1, 1, 1, 1, 1]  In Plane mosaic
        rotation and plane vectors (anisotropic),
3 mosaic_A, mosaic_B, A_h, A_k, A_l, B_h, B_k, B_l. Puts the crystal in
4 the in-plane mosaic state. Vectors A and B define plane in which
5 the crystal roation is defined, and mosaic_A, mosaic_B, denotes the
6 resp. mosaicities (gaussian RMS) with respect to the two
   reflections chosen by A and B (Miller indices).
   <0>:
8
9
   <0>:
10
   <0>:
11
   < 0 > :
12
   < 0 >:
13
   <0>:
14 < reflections = 0>: [string] File name containing structure factors of
       reflections. Use
15 empty ("") or NULL for incoherent scattering only
16 | <geometry=0>: [str] Name of an Object File Format (OFF) file for complex
       geometry.
17 The OFF file may be generated from XYZ coordinates using qhull/powercrust
18 < material = 0>:
19 < xwidth=0>: [m] Width of crystal
20 < yheight = 0>: [m] Height of crystal
21 < zdepth=0>: [m] Depth of crystal (no extinction simulated)
22 < radius=0>: [m] Outer radius of sample in (x,z) plane
23 < delta_d_d=1e-4>: [1] Lattice spacing variance, gaussian RMS
24 < mosaic = -1>: [arc minutes] Crystal mosaic (isotropic), gaussian RMS. Puts
       the crystal in the
   isotropic \ mosaic \ model \ state \,, \ thus \ disregarding \ other \ mosaicity \ parameters \,.
25
  <mosaic_a=-1>: [arc minutes] Horizontal (rotation around lattice vector a)
26
       mosaic (anisotropic), gaussian RMS.
27
   Put the crystal in the anisotropic crystal vector state. I.e. model
       mosaicity
   through rotation around the crystal lattice vectors. Has precedence over
   in-plane mosaic model.
30 < mosaic_b=-1>: [arc minutes] Vertical (rotation around lattice vector b)
       mosaic (anisotropic),
31
   gaussian RMS.
32 < mosaic_c=-1>: [arc minutes] Out-of-plane (Rotation around lattice vector c
       ) mosaic (anisotropic),
33 gaussian RMS
34 < recip_cell=0>: [1] Choice of direct/reciprocal (0/1) unit cell definition
35 \mid \text{cbarns=0} >: [1] \text{ Flag to indicate if } \mid \text{F} \mid ^2 \text{ from 'reflections' is in barns or}
36 barns=1 for laz, barns=0 for lau type files
37 < ax = 0 > : [-]
38 <ay=0>: [AA or AA^-1] Coordinates of first (direct/recip) unit cell vector
39 | < az = 0 > : [-]
40 | < bx = 0 > : [-]
```

```
41 | <br/>by=0>: [AA or AA^-1] Coordinates of second (direct/recip) unit cell vector
42 | < bz = 0 > : [-]
43 | < cx = 0 > : [-]
44 | <cy=0>: [AA or AA^-1] Coordinates of third (direct/recip) unit cell vector
45 | < cz = 0 > : [-]
46 <p_transmit=-1>: [1] Monte Carlo probability for photons to be transmitted
      without any scattering.
47 Used to improve statistics of the scattering pattern wrt. the attenuated,
      direct beam.
48 It is typically beneficial to leave this parameter at a value << 1. Use the
       value -1 to
49
  deactivate any stratified sampling and partition probabilities fully in
      accordance with the material cross-sections.
50 < sigma_abs=0>: [barns] absorption cross-section per unit cell at 2200 m/s
51 < sigma_inc=0>: [barns] incoherent scattering cross-section per unit cell
52 < aa = 0 > : [deg].
53 <bb=0>: [deg] unit cell angles alpha, beta and gamma. Then uses norms of
54 vectors a,b and c as lattice parameters
55 <cc=0>: [deg]
56 <order=0>: [1] limit multiple scattering up to given order
57 (0: all, 1: first, 2: second, ...)
```

Documentation Pending

7.6. Molecule_2state: Excitable time-dependent sample model

Input Parameters for component Molecule_2state from samples

```
1 | < Parameter = value >, [Unit], Description
 2 | <form_factors="FormFactors.txt">: [ ] File from which to read atomic form
       factors. Defualt amounts to use the one shipped with McXtrace.
3 < \text{state\_0\_file=NULL}>: [ ]  Isotropic scattering factors (parameterized by q),
       or atom positions are specified for state 0.
  <\!\!\text{state\_1\_file=\!NULL}\!\!>: [ \ ] \ Isotropic \ scattering \ factors \ (parameterized \ by \ q) \,,
       or atom positions are specified for state 1.
   <nq=512>: [ ] Number of q-bins if F is to be computed from atom positions (
      Debye formalism).
  <material_datafile="Be.txt">: [ ] Where to read f1 and f2 factors from in
      order to handle absorption.
7 | <delta_t=100e-9>: [s] Delay between the exciting event t=0. delay is
      negative, i.e. delta_t > 0 means the exciting event happens before t=0.
8 < excitation_yield = 0.2>: [ ] Mean fraction of molecules that get excited.
9 < t_relax = 100e - 9>: [s] Mean relaxation time (into state 0) of excited
  <initial_state=0>: [ ] Which state is Molecule_2state in for t<delta_t?</pre>
       Useful for modelling something that changes state slowly.
11 < psimin=0>: [rad] Minimum scattering angle off the optical axis.
12 | simax=M_PI_2>: [rad] Maximum scattering angle off the optical axis.
13 < etamin = -M.PI>: [rad] Minimum scattering angle around the optical axis.
14 < etamax=M_PI>: [rad] Maximum scattering angle around the optical axis.
15 < radius = 0>: [m] Radius of cylindrical of spherical sample.
16 < yheight = 0>: [m] Height of rectangular or cylindrical sample.
17 < xwidth=0>: [m] Width of rectangular sample.
18 < zdepth=0>: [m] Depth (thickness) of rectangular sample.
19 < concentration = 1>: [m] Concentration or packing factor of sample.
20 cp_transmit = 0.1 >: [m] Fraction of statistics devoted to sample direct (
      unscattered) beam.
  <q_parametric=0>: [ ] 0: Assume that datafiles contains atom positions. 1:
       datafiles contains F vs. q data.
   <Emax=80>: [ ] Maximal energy for which scattering factors are computed.
      Must be larger than the maximal impinging energy.
```

Further Documentaion pending

8. Monitors and detectors

In real scattering experiments, detectors and monitors play quite different roles. One wants the detectors to be as efficient as possible, counting all photons (absorbing them in the process), while the monitors measure the intensity of the incoming beam, and must as such be almost transparent, interacting only with (roughly) 0.1-1% of the photons passing by. In computer simulations, it is of course possible to detect every xray without absorbing it or disturbing any of its parameters. Hence, the two components have very similar functions in the simulations, and we do not distinguish between them. For simplicity, they are from here on just called **monitors**.

Another important difference between computer simulations and real experiments is that one may allow the monitor to be sensitive to any xray property, as *e.g.* direction, energy, and divergence, in addition to what is found in real-world detectors (space and time). One may, in fact, let the monitor record correlations between these properties.

When a monitor detects a xray, a number counting variable is incremented: $n_i = n_{i-1} + 1$. In addition, the photon weight p_i is added to the weight counting variable: $I_i = I_{i-1} + p_i$, and the second moment of the weight is updated: $M_{2,i} = M_{2,i-1} + p_i^2$. As also discussed chapter 2, after a simulation of N rays the detected intensity (in units of photonts/sec.) is I_N , while the estimated errorbar is $\sqrt{M_{2,N}^2}$.

Several different monitor components have been developed for McXtrace, but we have decided to support only the most important ones. One example of the monitors we have omitted is the single monitor, **Monitor**, that measures just one number (with errorbars) per simulation. This effect is mirrored by any of the 1- or 2-dimensional components we support, e.g. the PSD_monitor. In case additional functionality of monitors is required, a few lines of code in existing monitors can easily be modified.

Another solution is the "Swiss army knife" of monitors, **Monitor_nD**, that can handle almost any simulation requirement, but may prove challenging for inexperienced users or users who like to make their own modifications.

8.1. Monitor: Simple intensity monitor

Input Parameters for component Monitor from monitors

The **Monitor** component is a simple photon counter that merely detects the integrated intensity in an aperture xwidth by yheight m². It does not write a separate datafile, but reports the I, I_err, N -signals to the console. The signals represent intensity, error estimate, and number of statistical events (photon rays). If a scan is performed, the intensity recorded by **Monitor** will be written to the scan datafile.

8.2. E_{_}monitor: The energy-sensitive monitor

Input Parameters for component E_monitor from monitors

The component **E_monitor** is sensitive to the xray energy, which in binned in nE bins between E_{\min} and E_{\max} (in keV).

The output parameters from E_monitor are the total counts, and a file with 1-dimensional data vs. E, similar to TOF_monitor.

8.3. L_monitor: The wavelength sensitive monitor

Input Parameters for component L_monitor from monitors

```
| Parameter = value>, [Unit], Description
| NL=20>: [1] Number of wavelength channels
```

```
3 <filename=0>: [1] Name of file in which to store the detector image
4 <xmin=-0.05>: [m] Lower x bound of detector opening
5 <xmax=0.05>: [m] Upper x bound of detector opening
6 <ymin=-0.05>: [m] Lower y bound of detector opening
7 <ymax=0.05>: [m] Upper y bound of detector opening
8 <xwidth=0>: [m] Width of detector. Overrides xmin,xmax.
9 <yheight=0>: [m] Height of detector. Overrides ymin,ymax.
10 <Lmin>: [AA] Minimum wavelength to detect
11 <Lmax>: [AA] Maximum wavelength to detect
12 <restore_xray=0>: [1] If set, the monitor does not influence the xray state
13 <nowritefile=0>: [1] If set, monitor will skip writing to disk
```

The component **L_monitor** is very similar to **E_monitor**. This component is just sensitive to the xray wavelength. The wavelength spectrum is output in a one-dimensional histogram. between λ_{\min} and λ_{\max} (measured in Å).

As for the two other 1-dimensional monitors, this component outputs the total counts and a file with the histogram.

8.4. PSD_monitor: The PSD monitor

Input Parameters for component PSD_monitor from monitors

```
Parameter = value>, [Unit], Description

(filename=0>: [1] Name of file in which to store the detector image.

(xmin=-0.05>: [m] Lower x bound of detector opening

(xmax=0.05>: [m] Upper x bound of detector opening

(ymin=-0.05>: [m] Lower y bound of detector opening

(xmax=0.05>: [m] Upper y bound of detector opening

(xmidth=0>: [m] Width of detector. Overrides xmin,xmax.

(yheight=0>: [m] Height of detector. Overrides ymin,ymax.

(radius=0>: [m] Radius of circular detetor.

(restore_xray=1>: [1] If set, the monitor does not influence the xray state

(nowritefile=0>: [1] If set, monitor will skip writing to disk

(xx=90>: [1] Number of pixel columns.

(xy=90>: [1] Number of pixel rows.

(xr=0>: [1] Number of radial pixels.
```

The component PSD_monitor resembles other monitors, e.g. Monitor, and also propagates the photon ray to the detector surface in the (x, y)-plane, where the detector window is set by the x and y input coordinates. The PSD monitor is sensitive to the arrival position of the of the photon ray. The rectangular monitor window, given by the x and y limits is divided into $n_x \times n_y$ pixels.

The output from PSD_monitor is the integrated counts, n, I, M_2 , as well as three twodimensional arrays of counts: $n(x, y), I(x, y), M_2(x, y)$. The arrays are written to a file, filename, and can be read e.g. by the tool **mxplot**, see the system manual.

8.5. PSD_monitor_coh: The coherent PSD monitor

Input Parameters for component PSD_monitor_coh from monitors

```
1 | < Parameter = value >, [Unit], Description
2 | < nx = 90 > : []
                Number of pixel columns.
3 < ny = 90 > : [] Number of pixel rows.
  <filename=0>: [m] Name of file in which to store the detector images - the
      suffixes .abs and .arg will be added.
  <restore_xray=0>: [ ] If set, the monitor does not influence the xray state
  <xmin=-0.05>: [m] Lower x bound of detector opening.
6
7
  < xmax = 0.05 >: [m] Upper x bound of detector opening.
  <ymin=-0.05>: [m] Lower y bound of detector opening.
9 <ymax=0.05>: [m] Upper y bound of detector opening.
10 < xwidth=0>: [m] Width of detector. Overrides xmin, xmax.
11 < yheight=0>: [m] Height of detector. Overrides ymin, ymax.
12 < ratio=1>: [ ] ratio between pixel area and effective pixel area.
13 < centering = 1>: [ ] Treat all rays as if they hit the center of the pixel.
14 < nowritefile=0>: [1] If set, monitor will skip writing to disk
```

The component PSD_monitor_coh closely resembles its parent the **PSD_monitor**, and also propagates the photon ray to the detector surface in the (x, y)-plane, where the detector window is set by the x and y input coordinates. The coherent PSD monitor, though, is also sensitive considers the phase of the photon ray and emits not one but two datafiles, where one (filename.abs) represents the coherent intensity and the other (filename.arg) the phase of the collected beam.

The first output file (.abs) from PSD_monitor_coh is the integrated counts, n, I, M_2 , as well as three two-dimensional arrays of counts: $n(x, y), I(x, y), M_2(x, y)$. These arrays are written to a file, filename.abs, and can be read e.g. by the tool mxplot, see the system manual. Likewise, the second output file (.arg) contains the integrated phases, and are written to the file filename.arg

8.6. PSD_monitor_4PI: A $4\pi \operatorname{sr}$ spherical monitor.

Input Parameters for component PSD_monitor_4PI from monitors

PSD_monitor_4PI is an unphysical idealized component. It takes on the shape of a complete sphere with a given *radius*, records the longitude and latitude of photon rays as the pass through the sphere. The sphere is binned in constant spaced bins in longitude

and latitude. This has the consequence that the image will be distorted when plotted onto a plane, much like projections of a world map.

The output from PSD_monitor is the integrated counts, n, I, M_2 , as well as three twodimensional arrays of counts: $n(x, y), I(x, y), M_2(x, y)$. The arrays are written to a file, filename, and can be read e.g. by the tool **mxplot**, see the system manual.

8.7. EPSD_monitor: Energy-selective PSD monitor

Input Parameters for component EPSD_monitor from monitors

```
1 | < Parameter = value >, [Unit], Description
2 < nx = 90>: [m] Number of pixel columns
3 < ny = 90 >: [m] Number of pixel rows
4 | <filename=0>: [1] Name of file in which to store the detector image.
5 | <restore_xray=0>: [1] If set, the monitor does not influence the xray state
6
  <xmin=-0.05>: [m] Lower x bound of detector opening
  <xmax=0.05>: [m] Upper x bound of detector opening
7
  <ymin=-0.05>: [m] Lower y bound of detector opening
  <ymax=0.05>: [m] Upper y bound of detector opening
  <xwidth=0>: [m] Width of detector. Overrides xmin,xmax.
   <yheight=0>: [m] Height of detector. Overrides ymin,ymax.
12
  <Emax=0>: [keV] Upper bound of energy interval
13 <Emin=0>: [keV] Lower bound of energy interval
14 <nE=1>: [m] Number of energy bins
15 < nowritefile = 0>: [1] If set, monitor will skip writing to disk
```

The **EPSD_monitor** closely resembles **PSD_monitor** with the difference that this component may be given an energy interval: $[E_{min}, E_{max}]$, in which it is sensitive. Photon rays falling within this interval are detected, those outside are ignored.

The output from **EPSD_monitor** is the integrated counts, n, I, M_2 , as well as three two-dimensional arrays of counts: $n(x, y), I(x, y), M_2(x, y)$. The arrays are written to a file, filename, and can be read e.g. by the tool **mxplot**, see the system manual.

8.8. W_psd_monitor: A power vs. position monitor

Input Parameters for component W_psd_monitor from monitors

11 < yheight=0>: [m] Height of detector. Overrides ymin, ymax.

doc. pend.

8.9. Monitor_nD: A general Monitor for 0D/1D/2D records

Input Parameters for component Monitor_nD from monitors

```
1 < Parameter = value >, [Unit], Description
2 | <user1=FLT_MAX>: [variable] Variable assigned to User1
3 < user2=FLT_MAX>: [variable] Variable assigned to User2
4 | <user3=FLT_MAX>: [variable] Variable assigned to User3
5 < xwidth = 0>: [m] Width of detector.
6 < yheight=0>: [m] Height of detector.
7
  <zdepth=0>: [m] Thickness of detector (z).
  <xmin=0>: [m] Lower x bound of opening
9
  <xmax=0>: [m] Upper x bound of opening
10
  <ymin=0>: [m] Lower y bound of opening
11
  <ymax=0>: [m] Upper y bound of opening
12
  <zmin=0>: [m] Lower z bound of opening
13
  <zmax=0>: [m] Upper z bound of opening
  <bins=0>: [1] Number of bins to force for all variables. Use 'bins' keyword
14
       in 'options' for heterogeneous bins
  <min=-1e40>: [u] Minimum range value to force for all variables. Use 'min'
15
      or 'limits' keyword in 'options' for other limits
  <max=1e40>: [u] Maximum range value to force for all variables. Use 'max'
16
      or 'limits' keyword in 'options' for other limits
  <restore_xray=0>: [0|1] If set, the monitor does not influence the photon
17
      state. Equivalent to setting the 'parallel' option.
18 < radius = 0 >: [m] Radius of sphere / banana shape monitor
  <options="NULL">: [str] String that specifies the configuration of the
      monitor. The general syntax is "[x] options..." (see <b>Descr.</b>).
  <filename="NULL">: [str] Output file name (overrides file=XX option).
  <geometry="NULL">: [str] Name of an OFF file to specify a complex geometry
      detector
22 | <username1="NULL">: [str] Name assigned to User1
23 | <username2="NULL">: [str] Name assigned to User2
24 | <username3="NULL">: [str] Name assigned to User3
25 < nowritefile=0>: [1] If set, monitor will skip writing to disk
```

The component **Monitor_nD** is a general Monitor that may output any set of physical parameters regarding the passing photons. The generated files are either a set of 1D signals ([Intensity] vs. [Variable]), or a single 2D signal ([Intensity] vs. [Variable 1] vs. [Variable 2]), and possibly a simple long list of selected physical parameters for each photon ray.

The input parameters for **Monitor_nD** are its dimensions *xmin*, *xmax*, *ymin*, *ymax* (in metres) and an *options* string describing what to detect, and what to do with the signals, in clear language. The *xwidth*, *yheight*, *zdepth* may also be used to enter dimensions.

Eventhough the possibilities of Monitor_nD are numerous, its usage remains as simple as possible, specially in the *options* parameter, which 'understands' normal language. The formatting of the *options* parameter is free, as long as it contains some specific keywords, that can be sometimes followed by values. The *no* or *not* option modifier will revert next option. The *all* option can also affect a set of monitor configuration

parameters (see below).

As the usage of this component enables to monitor virtually anything, and thus the combinations of options and parameters is infinite, we shall only present the most basic configuration. The reader should refer to the on-line component help, using e.g. mcdoc Monitor_nD.comp.

8.9.1. The Monitor_nD geometry

The monitor shape can be selected among seven geometries:

- 1. (square) The default geometry is flat rectangular in (xy) plane with dimensions $x_{\min}, x_{\max}, y_{\min}, y_{\max}$, or x_{width}, y_{height} .
- 2. (box) A rectangular box with dimensions x_{width} , y_{height} , z_{depth} .
- 3. (disk) When choosing this geometry, the detector is a flat disk in (xy) plane. The radius is then

radius =
$$\max(\text{abs} [x_{\min}, x_{\max}, y_{\min}, y_{\max}, x_{width}/2, y_{height}/2]).$$
 (8.1)

- 4. (sphere) The detector is a sphere with the same radius as for the disk geometry.
- 5. (cylinder) The detector is a cylinder with revolution axis along y (vertical). The radius in (xz) plane is

$$radius = \max(abs [x_{min}, x_{max}, x_{width}/2]), \tag{8.2}$$

and the height along y is

$$height = |y_{max} - y_{max}| or y_{height}. \tag{8.3}$$

- 6. (banana) The same as the cylinder, but without the top/bottom caps, and on a restricted angular range. The angular range is specified using a theta variable limit specification in the options.
- 7. (previous) The detector has the shape of the previous component. This may be a surface or a volume. In this case, the photon is detected on the previous component, and there is no photon propagation.

By default, the monitor is flat, rectangular. Of course, you can choose the orientation of the **Monitor_nD** in the instrument description file with the usual ROTATED modifier.

For the box, sphere and cylinder, the outgoing photons are monitored by default, but you can choose to monitor incoming photons with the incoming option.

At last, the *slit* or *absorb* option will ask the component to absorb the photons that do not intersect the monitor. The *exclusive* option word removes photons which are similarly outside the monitor limits (that may be other than geometrical).

The parallel option keyword is of common use in the case where the **Monitor_nD** is superposed with other components. It ensures that photons are detected independently of other geometrical constrains. This is generally the case when you need e.g. to place more than one monitor at the same place.

8.9.2. The photon parameters that can be monitored

There are many different variables that can be monitored at the same time and position. Some can have more than one name (e.g. energy or omega).

```
kx ky kz k wavevector
                                              usually axis are
                               [Angs-1] (
 2
                                [m/s]
                                              x=horz., y=vert., z=on axis)
       vx vy vz v
3
                                [m]
                                         Distance, Position
       x y z
 4
       kxy vxy xy radius
                                [m]
                                         Radial wavevector, velocity and position
5
       t time
                                s
                                         Time of Flight
6
       energy omega
                                meV]
7
       lambda wavelength
                                Angs]
8
       p intensity flux
                                n/s] or [n/cm^2/s]
9
       ncounts
                                1
10
                                1]
                                         Spin
       sx sy sz
       vdiv ydiv dy
11
                                deg]
                                         vertical divergence (y)
12
       hdiv divergence xdiv
                                         horizontal divergence (x)
                                deg]
13
                                         divergence from direction
       angle
                                deg]
                               [\deg]
14
       theta longitude
                                         longitude (x/z) [for sphere and cylinder
15
              lattitude
                                         lattitude (y/z) [for sphere and cylinder
       phi
                               [deg]
```

as well as four other special variables

To tell the component what you want to monitor, just add the variable names in the *options* parameter. The data will be sorted into *bins* cells (default is 20), between some default *limits*, that can also be set by user. The *auto* option will automatically determine what limits should be used to have a good sampling of signals.

8.9.3. Important options

Each monitoring records the flux (sum of weights p) versus the given variables, except if the signal= $\langle variable \rangle$ word is used in the options.

The auto option is probably the most useful one: it asks the monitor to determine automatically the best limits for each variable, in order to obtain the most significant monitored histogram. This option should preced each variable, or be located after all variables in which case they are all affected. On the other hand, one may manually set the limits with the limits=[min max] option. If no limits are set monitor_nd uses predefined limits that usually make sense for most x-ray scattering simulations. Example: the default upper energy limit is 100 meV, but may be changed with an options string like options="energy limits 0 200". Note that the limits also apply in list mode (see below).

The log and abs options should be positioned before each variable to specify logarithmic binning and absolute value respectively.

The *borders* option will monitor variables that are outside the limits. These values are then accumulated on the 'borders' of the signal.

8.9.4. The output files

By default, the file names will be the component name, followed by a time stamp and automatic extensions showing what was monitored (such as MyMonitor.x). You can also set the filename in *options* with the *file* keyword followed by the file name that you want. The extension will then be added if the name does not contain a dot (.). Finally, the *filename* parameter may also be used.

The output files format are standard 1D or 2D McXtrace detector files. The *no file* option will *unactivate* monitor, and make it a single 0D monitor detecting integrated flux and counts. The *verbose* option will display the nature of the monitor, and the names of the generated files.

The 2D output

When you ask the **Monitor**_**nD** to monitor only two variables (e.g. options = "x y"), a single 2D file of intensity versus these two correlated variables will be created.

The 1D output

The **Monitor_nD** can produce a set of 1D files, one for each monitored variable, when using 1 or more than 2 variables, or when specifying the *multiple* keyword option.

The List output

The **Monitor_nD** can additionally produce a *list* of variable values for photons that pass into the monitor. This feature is additive to the 1D or 2D output. By default only 1000 events will be recorded in the file, but you can specify for instance "list 3000 photons" or "list all photons". This last option may require a lot of memory and generate huge files. Note that the limits to the measured parameters also apply in this mode. To exemplify, a monitor_nd instance with the option string "list all k" will only record those photons which have a below 2000 AA⁻¹, whereas an instance with the option string "list all kx ky kz 0 2000" will record all photons with $|k_x, k_y| < 2000$ AA⁻¹ and $0 < v_z < 2000$ AA⁻¹. Thus, in this latter case, any neutron travelling in the negative z-direction will be disregarded.

8.9.5. Monitor equivalences

In the following table 8.1, we show how the Monitor_nD may substitute any other McX-trace monitor.

McXtrace monitor	Monitor_nD equivalent
E_monitor	options="energy bins=nchan limits= $[E_{min}E_{max}]$ "
EPSD_monitor	options="energy bins= n_E limits= $[E_{min}E_{max}]$, x bins= nx "
$L_{-}monitor$	$xmin=x_{min} xmax=x_{max}$ $options=$ "lambda bins= nh limits= $[-\lambda_{max}/2\lambda_{max}/2]$ " file- name=file
Monitor	options="unactivate"
PSD_monitor_4PI	options="theta y, sphere"
PSD_monitor	$options = "x bins = nx, y bins = ny" xmin = x_{min} xmax = x_{max}$
	$ymin=y_{min} ymax=y_{max} filename=file$

Table 8.1.: Using Monitor_nD in place of other components. All limits specifications may be advantageously replaced by an *auto* word preceding each monitored variable. Not all file and dimension specifications are indicated (e.g. filename, xmin, xmax, ymin, ymax).

8.9.6. Usage examples

```
COMPONENT MyMonitor = Monitor\_nD(

xmin = -0.1, xmax = 0.1,

ymin = -0.1, ymax = 0.1,

options = "energy auto limits")
```

will monitor the photon energy in a single 1D file (a kind of E_monitor)

- options = "banana, theta limits=[10,130], bins=120, y bins=30" is a theta/height banana detector.
- options = "banana, theta limits=[10,130], auto time" is a theta/time-of-flight banana detector.
- options="x bins=30 limits= $[-0.05 \ 0.05]$; y" will set the monitor to look at x and y. For y, default bins (20) and limits values (monitor dimensions) are used.
- options="x y, auto, all bins=30" will determine itself the required limits for x and y.
- options="multiple x bins=30, y limits=[-0.05 0.05], all auto" will monitor the photon x and y in two 1D files.
- options="x y z kx ky kz, all auto" will monitor each of these variables in six 1D files.
- options="x y z kx ky kz, list all, all auto" will monitor all these photons' variables in one long list, one row per photon event.

- options="multiple x y z kx ky kz, and list 2000, all auto" will monitor all the photons' variables in one list of 2000 events and in six 1D files.
- options="signal=energy, x y" is a PSD monitor recording the mean energy of the beam as a function of x and y.

8.9.7. Monitoring user variables

There are two ways to monitor any quantity with Monitor_nD. This may be e.g. the number of reflections in a mirror system, or the wavevector and energy transfer at a sample. The only requirement is to define the user1 (and optionally user2,user3) variables of a given Monitor_nD instance.

Directly setting the user variables (simple)

The first method uses the user1 and username1 component parameters to directly transfer the value and label, such as in the following example:

```
TRACE
(...)
COMPONENT UserMonitor = Monitor\_nD(
user1 = log(t), username1="Log(time)",
options ="auto user1")
```

The values to assign to user2 and user3 must be global instrument variables, or a component output variable as in user1=MC_GETPAR(some_comp, outpar). Similarly, the user2,user3 and username2,username3 parameters may be used to control the second and third user variable, to produce eventually 2D/3D user variable correlation data and custom event lists.

Setting indirectly the user variables (only for professionals)

It is possible to control the user variables of a given Monitor_nD instance anywhere in the instrument description. This method requires more coding, but has the advantage that a variable may be defined to store the result of a computation locally, and then transfer it into the UserMonitor, all fitting in an EXTEND block.

This is performed in a 4-step process:

1. Declare that you intend to monitor user variables in a Monitor_nD instance (defined in TRACE):

```
DECLARE

%{ (...)

%include "monitor_nd-lib"

MONND\_DECLARE(UserMonitor); // will monitor custom things in UserMonitor

%}
```

2. Initialize the label of the user variable (optional):

The value '1' could be '2' or '3' for the user2, user3 variable.

3. Set the user variable value in a TRACE component EXTEND block:

```
TRACE
(...)
COMPONENT blah = blah\_comp(...)
EXTEND

%{    // attach a value to user1 in UserMonitor, could be much more complex here.
MONND\_USER\_VALUE(UserMonitor, 1, log(t));
%}
(...)
```

4. Tell the Monitor_nD instance to record user variables:

```
TRACE
(...)
COMPONENT UserMonitor = Monitor\_nD(options="auto user1")
4
(...)
```

Setting the user variable values may either make use of the photon parameters (x,y,z, vx,vy,vz, phi, t, Ex,Ey,Ez, p), access the internal variables of the component that sets the user variables (in this example, those from the blah instance), access any component OUTPUT parameter using the MC_GETPAR C macro(see chapter A), or simply use a global instrument variable. Instrument parameters can not be used directly.

8.10. PreMonitor_nD: Store photon rays for possible later detection.

Input Parameters for component PreMonitor_nD from monitors

```
Parameter = value>, [Unit], Description
```

The first immediate usage of the **Monitor_nD** component is when one requires to identify cross-correlations between some photon parameters, e.g. position and divergence (aka phase-space diagram). This latter monitor would be merely obtained with:

```
options="x dx, auto", bins=30
```

This example records the correlation between position and divergence of photons at a given instrument location.

It is also possible to search for cross-correlation between two part of the instrument simulation. One example is the acceptance phase-diagram, which shows the photon caracteristics at the input required to reach the end of the simulation. This *spatial* correlation may be revealed using the **PreMonitor_nD** component. This latter stores the photon parameters at a given instrument location, to be used at an other **Monitor_nD** location for monitoring.

The only parameter of **PreMonitor_nD** is the name of the associated Monitor_nD instance, which should use the **premonitor** option, as in the following example:

```
COMPONENT CorrelationLocation = PreMonitor_nD(comp = CorrelationMonitor)
AT (...)
    (... e.g. a bunch of optics )

COMPONENT CorrelationMonitor = Monitor_nD(
    options="x dx, auto, all bins=30, premonitor")
AT (...)
```

which performs the same monitoring as the previous example, but with a spatial correlation constrain. Indeed, it records the position vs the divergence of photons at the correlation location, but only if they reach the monitoring position. All usual **Monitor_nD** variables may be used, except the user variables. The latter may be defined as described in section 8.9.7 in an EXTEND block.

9. Special-purpose components

The chapter deals with components that are not easily included in any of the other chapters because of their special nature, but which are still part of the McXtrace system.

One part of these components deals with splitting simulations into two (or more) stages. For example, the front end of a beamline is often not changed much, and a long simulation of photon rays "surviving" through initial optics could be reused for several simulations of the beamline back-end, speeding up the simulations by (typically) one or two orders of magnitude. The components for doing this trick is **Virtual_input** and **Virtual_output**, which stores and reads photon rays, respectively.

For integration with SHADOW [Rio+11] the components **Shadow_inpt** and **Shadow_output** have been written.

Progress_bar is a simulation utility that displays the simulation status, but assumes the form of a component.

9.1. Progress_bar: Dynamic information output

Input Parameters for component Progress_bar from misc

The component **Progress_bar** is a special version of an Arm and has no effect on the x-ray. Only one instance of progress bar is allowed - usually it is the very first component of a beamline simulation.

If percent > 0 (default is 10%) the porgress bar prints an update when approximately percent % of the total number of rays have been simulated. If minutes > 0 (overrides percen) this update happens approx. every minutes min. If $flag_save \neq 0$, the update process also includes saving the status of all monitors to disk. This is useful if something is likely to interrupt the progress of long simulations of avoid loss of data. profile may be set to save the simulation profile to a filename other than the name of the instrument (the default).

9.2. Virtual_output: Saving the first part of a split simulation

Input Parameters for component Virtual_output from misc

```
1 < Parameter = value >, [Unit], Description
```

The component **Virtual_output** stores the photon ray parameters at the end of the first part of a split simulation. The idea is to let the next part of the split simulation be performed by another instrument file, which reads the stored photon ray parameters by the component **Virtual_input**.

All photon ray parameters are saved to the output file, which is by default of "text" type, but can also assume the binary formats "float" or "double". The storing of photon rays continues until the specified number of simulations have been performed.

buffer-size may be used to limit the size of the output file, but absolute intentities are then likely to be wrong. Exept when using MPI, we recommend to use the default value of zero, saving all photon rays. The size of the file is then controlled indirectly with the general ncounts parameter.

9.3. Virtual_input: Starting the second part of a split simulation

Input Parameters for component Virtual_input from misc

```
1 < Parameter = value >, [Unit], Description
```

The component **Virtual_input** resumes a split simulation where the first part has been performed by another instrument and the photon ray parameters have been stored by the component **Virtual_output**.

All photon ray parameters are read from the input file, which is by default of "text" type, but can also assume the binary formats "float" and "double". Reading of photon rays continues until the specified number of rays have been simulated or the file has been exhausted. If desirable, the input file can be reused a number of times, determined by the optional parameter repeat-count. This is only useful if the present simulation makes use of MC choices, otherwise the same outcome will result for each repetition of the simulation (see chapter 2).

Care should be taken when dealing with absolute intensities, which will be correct only when the input file has been exhausted at least once.

The simulation ends with either the end of the repeated file counts, or with the normal end with *ncount* McXtrace simulation events. We recommend controlling the simulation on *repeat-count* by using a larger ncount value.

9.4. Shadow_input: Reading input from Shadow

Input Parameters for component Shadow_input from misc

Documentation pending

9.5. Shadow_output: Saving the photon rays for use with SHADOW

Input Parameters for component Shadow_output from misc

Documentation pending.

A. Libraries and conversion constants

The McXtrace Library contains a number of built-in functions and conversion constants which are useful when constructing components. These are stored in the share directory of the MCXTRACE library.

Within these functions, the 'Run-time' part is available for all component/instrument descriptions. The other parts are dynamic, that is they are not pre-loaded, but only imported once when a component requests it using the **%include** McXtrace keyword. For instance, within a component C code block, (usually SHARE or DECLARE):

```
%include "read_table-lib"
```

will include the 'read_table-lib.h' file, and the 'read_table-lib.c' (unless the --no-runtime option is used with mcxtrace). Similarly,

```
%include "read_table-lib.h"
```

will *only* include the 'read_table-lib.h'. The library embedding is done only once for all components (like the SHARE section). For an example of implementation, see **Res_monitor**.

In this Appendix, we present a short list of both each of the library contents and the run-time features.

A.1. Run-time calls and functions (mcxtrace-r)

Here we list a number of preprogrammed macros and functions which may ease the task of writing component and instrument definitions. By convention macros are in upper case whereas functions are in lower case.

A.1.1. Photon propagation

Propagation routines perform all necessary operations to transport x-rays from one point to an other. Except when using the special ALLOW_BACKPROP; call prior to executing any PROP_* propagation, the x-rays which have negative propagation lengths are removed automatically.

- **ABSORB**. This macro issues an order to the overall McXtrace simulator to interrupt the simulation of the current x-ray history and to start a new one.
- **PROP_ZO**. Propagates the x-ray to the z=0 plane, by adjusting (x,y,z), ϕ , and t accordingly from knowledge of the x-ray wavevector (kx,ky,kz). If the propagation length is negative, the x-ray is absorbed, except if a ALLOW_BACKPROP; preceeds it.

For components that are centered along the z-axis, use the _intersect functions to determine intersection time(s), and then a PROP_DL call.

- **PROP_X0**, **PROP_Y0**. These macros are analogous to PROP_Z0 except they propagate to the x = 0 and y = 0 planes respectively.
- **PROP_DL**(dl). Propagates the x-ray by the length dl, adjusting (x, y, z), ϕ , t accordingly, from knowledge of the x-ray wavevector.
- ALLOW_BACKPROP. Indicates that the *next* propagation routine will not remove the x-ray, even if negative propagation lengths are found. Subsequent propagations are not affected.
- SCATTER. This macro is used to denote a scattering event inside a component. It should be used to indicate that a component has interacted with the x-ray (e.g. scattered or detected). This does not affect the x-ray state (see, however, Beamstop), and it is mainly used by the MCDISPLAY section and the GROUP modifier. See also the SCATTERED variable (below).

A.1.2. Coordinate and component variable retrieval

- MC_GETPAR(comp, outpar). This may be used in e.g. the FINALLY section of an instrument definition to reference the parameters of a component.
- NAME_CURRENT_COMP gives the name of the current component as a string.
- POS_A_CURRENT_COMP gives the absolute position of the current component. A component of the vector is referred to as POS_A_CURRENT_COMP.i where i is x, y or z.
- ROT_A_CURRENT_COMP and ROT_R_CURRENT_COMP give the orientation of the current component as rotation matrices (absolute orientation and the orientation relative to the previous component, respectively). A component of a rotation matrix is referred to as ROT_A_CURRENT_COMP[m][n], where m and n are 0, 1, or 2 standing for x, y and z coordinates respectively.
- **POS_A_COMP**(comp) gives the absolute position of the component with the name comp. Note that comp is not given as a string. A component of the vector is referred to as POS_A_COMP(comp).i where i is x, y or z.
- ROT_A_COMP(comp) and ROT_R_COMP(comp) give the orientation of the component comp as rotation matrices (absolute orientation and the orientation relative to its previous component, respectively). Note that comp is not given as a string. A component of a rotation matrice is referred to as ROT_A_COMP(comp)[m][n], where m and n are 0, 1, or 2.

- **INDEX_CURRENT_COMP** is the number (index) of the current component (starting from 1).
- POS_A_COMP_INDEX(index) is the absolute position of component index. POS_A_COMP_INDEX (INDEX_CURRENT_COMP) is the same as POS_A_CURRENT_COMP. You may use POS_A_COMP_INDEX (INDEX_CURRENT_COMP+1) to make, for instance, your component access the position of the next component (this is usefull for automatic targeting). A component of the vector is referred to as POS_A_COMP_INDEX(index).i where i is x, y or z.
- POS_R_COMP_INDEX works the same as above, but with relative coordinates.
- STORE_XRAY(index, x, y, z, kx, ky, kz, phi, t, Ex, Ey, Ez, p) stores the current x-ray state in the trace-history table, in local coordinate system. index is usually INDEX_CURRENT_COMP. This is automatically done when entering each component of an instrument.
- RESTORE_XRAY(index, x, y, z, kx, ky, kz, phi, t, Ex, Ey, Ez, p) restores the x-ray state to the one at the input of the component index. To ignore a component effect, use RESTORE_XRAY (INDEX_CURRENT_COMP, x, y, z, kx, ky, kz, phi, Ex, Ey, Ez, p) at the end of its TRACE section, or in its EXTEND section. These x-ray states are in the local component coordinate systems.
- **SCATTERED** is a variable set to 0 when entering a component, which is incremented each time a SCATTER event occurs. This may be used in the EXTEND sections to determine whether the component interacted with the current x-ray.
- extend_list(n, &arr, &len, elemsize). Given an array arr with len elements each of size elemsize, make sure that the array is big enough to hold at least n elements, by extending arr and len if necessary. Typically used when reading a list of numbers from a data file when the length of the file is not known in advance.
- $\mathbf{mcset_ncount}(n)$. Sets the number of x-ray histories to simulate to n.
- mcget_ncount(). Returns the number of x-ray histories to simulate (usually set by option -n).
- mcget_run_num(). Returns the number of x-ray histories that have been simulated until now.

A.1.3. Coordinate transformations

- **coords_set**(x, y, z) returns a Coord structure (like POS_A_CURRENT_COMP) with x, y and z members.
- **coords_get**(P, &x, &y, &z) copies the x, y and z members of the Coord structure P into x, y, z variables.

- $\mathbf{coords_add}(a, b)$, $\mathbf{coords_sub}(a, b)$, $\mathbf{coords_neg}(a)$ enable to operate on coordinates, and return the resulting Coord structure.
- rot_set_rotation(Rotation t, ϕ_x , ϕ_y , ϕ_z) Get transformation matrix for rotation first ϕ_x around x axis, then ϕ_y around y, and last ϕ_z around z. t should be a 'Rotation' ([3][3] 'double' matrix).
- $rot_mul(Rotation\ t1,\ Rotation\ t2,\ Rotation\ t3)\ performs\ t3 = t1.t2.$
- $rot_copy(Rotation \ dest, \ Rotation \ src)$ performs dest = src for Rotation arrays.
- $rot_transpose(Rotation\ src,\ Rotation\ dest)$ performs $dest = src^t$.
- rot_apply(Rotation t, Coords a) returns a Coord structure which is t.a

A.1.4. Mathematical routines

- NORM(x, y, z). Normalizes the vector (x, y, z) to have length 1.
- scalar_prod $(a_x, a_y, a_z, b_x, b_y, b_z)$. Returns the scalar product of the two vectors (a_x, a_y, a_z) and (b_x, b_y, b_z) .
- **vec_prod**(& a_x ,& a_y ,& a_z , b_x , b_y , b_z , c_x , c_y , c_z). Sets (a_x, a_y, a_z) equal to the vector product $(b_x, b_y, b_z) \times (c_x, c_y, c_z)$.
- rotate(&x,&y,&z, v_x , v_y , v_z , φ , a_x , a_y , a_z). Set (x, y, z) to the result of rotating the vector (v_x, v_y, v_z) the angle φ (in radians) around the vector (a_x, a_y, a_z) .
- **normal_vec** (n_x, n_y, n_z, x, y, z) . Computes a unit vector (n_x, n_y, n_z) normal to the vector (x, y, z).*
- solve_2nd_order(* t_0 ,* t_1 , A, B, C). Solves the 2^{nd} order equation $At^2 + Bt + C = 0$ and puts the solutions in * t_0 and * t_1 . The smallest positive solution into pointer * t_0 . If t_1 =NULL it is ignored and the second solution is discarded.

A.1.5. Output from detectors

Details about using these functions are given in the McXtrace User Manual.

- **DETECTOR_OUT_0D(...)**. Used to output the results from a single detector. The name of the detector is output together with the simulated intensity and estimated statistical error. The output is produced in a format that can be read by McXtrace front-end programs.
- **DETECTOR_OUT_1D**(...). Used to output the results from a one-dimensional detector. Integrated intensities error etc. is also reported as for DETECTOR_OUT_0D.
- **DETECTOR_OUT_2D**(.....). Used to output the results from a two-dimentional detector. Integrated intensities error etc. is also reported as for DETECTOR_OUT_0D.

• mcinfo_simulation(FILE *f, mcformat, char *pre, char *name) is used to append the simulation parameters into file f (see for instance Res_monitor). Internal variable mcformat should be used as specified. Please contact the authors for further information.

A.1.6. Ray-geometry intersections

- inside_rectangle(x, y, xw, yh). Return 1 if $-xw/2 \le x \le xw/2$ AND $-yh/2 \le y \le yh/2$. Else return 0.
- box_intersect(& l_1 , & l_2 , x, y, z, k_x , k_y , k_z , d_x , d_y , d_z). Calculates the (0, 1, or 2) intersections between the x-ray path and a box of dimensions d_x , d_y , and d_z , centered at the origin for a x-ray with the parameters (x, y, z, k_x, k_y, k_z) . The intersection lengths are returned in the variables l_1 and l_2 , with $l_1 < l_2$. In the case of less than two intersections, t_1 (and possibly t_2) are set to zero. The function returns true if the x-ray intersects the box, false otherwise.
- cylinder_intersect(& l_1 , & l_2 , x, y, z, k_x , k_y , k_z , r, h). Similar to box_intersect, but using a cylinder of height h and radius r, centered at the origin.
- sphere_intersect(& l_1 , & l_2 , x, y, z, k_x , k_y , k_z , r). Similar to box_intersect, but using a sphere of radius r.
- ellipsoid_intersect(& l_1 , & l_2 , x, y, z, k_x , k_y , k_z , a,b,c,Q,). Similar to box_intersect, but using an ellipsoid with half-axis a,b,c oriented by the rotation matrix Q. If Q = I, a is along the x-axis, b along y and c along z

A.1.7. Random numbers

By default McXtrace uses the included Mersenne Twister[MN98] algorithm for generating pseudo random numbers.

- rand01(). Returns a random number distributed uniformly between 0 and 1.
- randnorm(). Returns a random number from a normal distribution centered around 0 and with $\sigma = 1$. The algorithm used to sample the normal distribution is explained in Ref. [Pre+86, ch.7].
- randpm1(). Returns a random number distributed uniformly between -1 and 1.
- randtriangle(). Returns a random number from a triangular distribution between -1 and 1.
- randvec_target_circle(& v_x , & v_y , & v_z , & $d\Omega$, aim $_x$, aim $_y$, aim $_z$, r_f). Generates a random vector (v_x , v_y , v_z), of the same length as (aim $_x$, aim $_y$, aim $_z$), which is targeted at a disk centered at (aim $_x$, aim $_y$, aim $_z$) with radius r_f (in meters), and perpendicular to the aim vector.. All directions that intersect the circle are

chosen with equal probability. The solid angle of the circle as seen from the position of the x-ray is returned in $d\Omega$. This routine was previously called **rand-vec_target_sphere** (which still works).

- randvec_target_rect_angular(& v_x , & v_y , & v_z , & $d\Omega$, aim $_x$, aim $_y$, aim $_z$, h, w, Rot) does the same as randvec_target_circle but targetting at a rectangle with angular dimensions h and w (in radians, not in degrees as other angles). The rotation matrix Rot is the coordinate system orientation in the absolute frame, usually ROT_A_CURRENT_COMP.
- randvec_target_rect(& v_x , & v_y , & v_z , & $d\Omega$, aim $_x$, aim $_y$, aim $_z$, height, width, Rot) is the same as randvec_target_rect_angular but height and width dimensions are given in meters. This function is useful to e.g. target at a guide entry window or analyzer blade.

A.2. Reading a data file into a vector/matrix (Table input, read_table-lib)

The read_table-lib library provides functionalities for reading text (and binary) data files. To use this library, add a %include "read_table-lib" in your component definition DECLARE or SHARE section. Tables are structures of type t_Table (see read_table-lib.h file for details):

```
/* t_Table structure (most important members) */
2
                          /* Use Table_Index(Table, i j) to extract [i,j]
      double *data;
          element */
3
                          /* number of rows */
      long
               rows;
      long
                          /* number of columns */
4
               columns;
                          /* the header with comments */
5
      char
              *header:
6
                          /* file name or title */
      char
              *filename;
                          /* minimum value of 1st column/vector */
      double
               \min_{\mathbf{x}}:
                          /* maximum value of 1st column/vector */
      double
               \max_{x};
```

Available functions to read a single vector/matrix are:

- Table_Init(&Table, rows, columns) returns an allocated Table structure. Use rows = columns = 0 not to allocate memory and return an empty table. Calls to Table_Init are optional, since initialization is being performed by other functions already.
- Table_Read(& Table, filename, block) reads numerical block number block (0 to catenate all) data from text file filename into Table, which is as well initialized in the process. The block number changes when the numerical data changes its size, or a comment is encoutered (lines starting by '#; %/'). If the data could not be read, then Table.data is NULL and Table.rows = 0. You may then try to read it using Table_Read_Offset_Binary. Return value is the number of elements read.

- Table_Read_Offset(&Table, filename, block, &offset, n_{rows}) does the same as Table_Read except that it starts at offset offset (0 means beginning of file) and reads n_{rows} lines (0 for all). The offset is returned as the final offset reached after reading the n_{rows} lines.
- Table_Read_Offset_Binary(&Table, filename, type, block, &offset, n_{rows} , $n_{columns}$) does the same as Table_Read_Offset, but also specifies the type of the file (may be "float" or "double"), the number n_{rows} of rows to read, each of them having $n_{columns}$ elements. No text header should be present in the file.
- Table_Rebin(&Table) rebins all Table rows with increasing, evenly spaced first column (index 0), e.g. before using Table_Value. Linear interpolation is performed for all other columns. The number of bins for the rebinned table is determined from the smallest first column step.
- Table_Info(Table) print information about the table Table.
- Table_Index(Table, m, n) reads the Table[m][n] element.
- Table_Value(Table, x, n) looks for the closest x value in the first column (index 0), and extracts in this row the n-th element (starting from 0). The first column is thus the 'x' axis for the data.
- Table_Free(&Table) free allocated memory blocks.
- Table_Value2d(Table, X, Y) Uses 2D linear interpolation on a Table, from (X,Y) coordinates and returns the corresponding value.

Available functions to read an array of vectors/matrices in a text file are:

- Table_Read_Array(File, &n) read and split file into as many blocks as necessary and return a t_Table array. Each block contains a single vector/matrix. This only works for text files. The number of blocks is put into n.
- Table_Free_Array(&Table) free the Table array.
- Table_Info_Array(&Table) display information about all data blocks.

The format of text files is free. Lines starting by '#; % /' characters are considered to be comments, and stored in *Table.header*. Data blocks are vectors and matrices. Block numbers are counted starting from 1, and changing when a comment is found, or the column number changes. For instance, the file 'MCXTRACE/data/Rh.txt' (Material data for Rhodium) looks like:

```
1 #Rh (Z 45)
2 #Atomic weight: A[r] 102.9055
3 #Nominal density: rho 1.2390E+01
4 # sigma[a](barns/atom) = [mu/rho](cm\^2 g\^-1) . 1.70879E+02
5 # E(eV) [mu/rho](cm\^2 g\^-1) = f[2](e atom\^-1) . 4.08922E+05
6 # 14 edges. Edge energies (keV):
```

```
7 #
8 #
                                                                3.14610E+00 L III
9 #
                   2.32199E+01 L I
                                         3.41190E+00 L II
          Κ
       3.00380E+00
                                         5.21000E-01 M III
10 #
                   6.27100E-01 M II
                                                                4.96200E-01 M IV
          M I
       3.11700E-01
                   3.07000E-01
                                NI
                                         8.10000E-02 N II
                                                                4.79000E-02 N III
  #
          M V
11
       4.79000E-02
12
  #
          N IV
                   2.50000E-03 N V
                                         2.50000E-03
13 #
14 #
         Relativistic correction estimate f[rel] (H82,3/5CL) = -4.0814E-01,
15 #
         -2.5440E-01 \text{ e atom}^-1
         Nuclear Thomson correction f[NT] = -1.0795E-02 \text{ e atom}^-1
16 #
17
  #
18 #
19 #Form Factors, Attenuation and Scattering Cross-sections
20 | \#Z = 45, E = 0.001 - 433 \text{ keV}
21 #
                                                                       [sigma/rho]
22 #
                                         f [2]
                                                       [mu/rho]
             [mu/rho]
                             [mu/rho][K]
                                               lambda
23
                                               Photoelectric Coh+inc
                                                                             Total
  #
24 #
                      e atom^-1
                                        e atom^-1
                                                       cm^2 g^-1
                                                                           cm^2 g^-1
                             cm^2 g^-1
             cm^2 g^-1
                                               1.8382E+05
25
   1.069000E-02
                  1.89417E+00
                                 4.8055E+00
                                                            1.1514E-04
                                                                          1.8382\mathrm{E}{+}05
       0.000E+00
                   1.160E+02
26
   1.142761E\!\!-\!\!02
                   2.09662\mathrm{E}{+00}
                                 5.1028E+00
                                               1.8260E+05
                                                            1.5865E-04
                                                                          1.8260E+05
       0.000E+00
                   1.085E+02
27
   1.221612E-02
                   2.32705E+00
                                 5.4019E+00
                                               1.8082E+05
                                                            2.1741 E\!-\!04
                                                                          1.8082E+05
       0.000E+00
                   1.015E+02
                                 5.6998E+00
                                               1.7848E+05
                                                            2.9628E-04
                                                                          1.7848E+05
28
   1.305903E-02
                   2.58575E+00
       0.000E+00
                   9.494E + 01
29
   1.396010E-02
                   2.87263E+00
                                 5.9931\mathrm{E}{+00}
                                               1.7555E+05
                                                            4.0158E-04
                                                                          1.7555 \mathrm{E}{+}05
       0.000E+00
                   8.881E+01
30
                                 6.2786E+00
                                               1.7204E+05
                                                            5.4136E-04
                                                                          1.7204\mathrm{E}{+}05
   1.492335E-02
                   3.18714E+00
       0.000E+00
                   8.308E+01
31
   1.595306E-02
                   3.52819E+00
                                 6.5531E+00
                                               1.6797E+05
                                                            7.2588E-04
                                                                          1.6797E+05
                   7.772E+01
       0.000E+00
   1.705382E-02
                   3.89415E+00
                                 6.8134E+00
                                               1.6337E+05
                                                            9.6809E - 04
                                                                          1.6337E+05
       0.000E+00
                   7.270E+01
33
     . . .
```

Binary files should be of type "float" (i.e. REAL*32) and "double" (i.e. REAL*64), and should *not* contain text header lines. These files are platform dependent (little or big endian).

The *filename* is first searched into the current directory (and all user additional locations specified using the ¬I option, see the 'Running McXtrace' chapter in the User Manual), and if not found, in the data sub-directory of the MCXTRACE library location. This way, you do not need to have local copies of the McXtrace Library Data files (see table 1.1).

A usage example for this library part may be:

```
t_Table Table;
1
                             // declare a t_Table structure
     char file[]="Rh.txt"; // a file name
2
3
     double x,y;
4
     Table_Read(&Table, file, 1); // initialize and read the first numerical
5
         block
                                       // display table informations
6
     Table_Info(Table);
7
     x \,=\, Table\_Index\left(\, Table \,,\, \, 2\,,5\right); \quad // \ \textit{read the 3rd row}\,, \ \textit{6th column element}
8
9
                                       // of the table. Indexes start at zero in C
     y = Table_Value (Table, 1.45,1); // look for value 1.45 in 1st column (x
10
         axis)
                                       // and extract 2nd column value of that row
11
     Table_Free(&Table);
                                        // free allocated memory for table
12
```

Additionally, if the block number (3rd) argument of **Table_Read** is 0, all blocks will be catenated. The **Table_Value** function assumes that the 'x' axis is the first column (index 0). Other functions are used the same way with a few additional parameters, e.g. specifying an offset for reading files, or reading binary data.

This other example for text files shows how to read many data blocks:

```
1
     t_Table *Table;
                            // declare a t_Table structure array
2
     long
              \mathrm{n}\,;
3
    double y;
    Table = Table_Read_Array("file.dat", &n); // initialize and read the all
        numerical block
                                        // display informations for all blocks (
6
    n = Table_Info_Array (Table);
        also returns n)
7
    y = Table\_Index(Table[0], 2,5); // read in 1st block the 3rd row, 6th
8
         column element
9
                                        // ONLY use Table[i] with i < n!
10
     Table_Free_Array (Table);
                                        // free allocated memory for Table
```

You may look into, for instance, the source files for **Lens_parab** or **Filter** for other implementation examples.

A.3. Constants for unit conversion etc.

The following predefined constants are useful for conversion between units

Name	Value	Conversion from	Conversion to
DEG2RAD	$2\pi/360$	Degrees	Radians
RAD2DEG	$360/(2\pi)$	Radians	Degrees
MIN2RAD	$2\pi/(360 \cdot 60)$	Minutes of arc	Radians
RAD2MIN	$(360 \cdot 60)/(2\pi)$	Radians	Minutes of arc
FWHM2RMS	$1/\sqrt{8\log(2)}$	Full width half maximum	Root mean square
	•		(standard deviation)
RMS2FWHM	$\sqrt{8\log(2)}$	Root mean square (stan-	Full width half maxi-
	•	dard deviation)	mum
MNEUTRON	$1.67492 \cdot 10^{-27} \text{ kg}$	Neutron mass, $m_{\rm n}$	
HBAR	$1.05459 \cdot 10^{-34} \text{ Js}$	Planck constant, \hbar	
PI	3.14159265	π	
CELE	1.602176487e-19	Elementary charge (C)	
$M_{-}C$	299792458	Speed of light in vacuum	
		(m/s)	
NA	6.02214179e23	Avogadro's number	
		(#atoms/g·mole)	
RE	2.8179402894e-5	Thomson scattering	
		length (AA)	
E2K	0.506773091264796	Wavenumber (1/AA)	Energy (keV)
K2E	1.97326972808327	Energy (keV)	Wavenumber (1/AA)

B. The McXtrace terminology

This is a short explanation of phrases and terms which have a specific meaning within McXtrace. We have tried to keep the list as short as possible running the calculated risk that the reader may occasionally miss an explanation. In this case, you are more than welcome to contact the McXtrace core team.

- **Arm** A generic McXtrace component which defines a frame of reference for other components.
- Component One unit (e.g. optical element) in an x-ray beamline. These are considered as Types of elements to be instantiated in an Instrument description.
- Component Instance A named Component (of a given Type) inserted in an Instrument description.
- **Definition parameter** An input parameter for a component. For example the radius of a sample component or the divergence of a collimator. Technically, a definition parameter is translated into a literal constant, which prevents it from being edited at runtime.
- **Input parameter** For a component, either a definition parameter or a setting parameter. These parameters are supplied by the user to define the characteristics of the particular instance of the component definition. For an instrument, a parameter that can be changed at simulation run-time.
- **Instrument** An assembly of McXtrace components defining an x-ray beamline.
- Kernel The McXtrace language definition and the associated compiler
- Output parameter An output parameter for a component. For example the counts in a monitor. An output parameter may be accessed from the instrument in which the component is used using MC_GETPAR.
- Run-time C code, contained in the files mcxtrace-r.c and mcxtrace-r.h included in the McXtrace distribution, that declare functions and variables used by the generated simulations.
- Setting parameter Similar to a definition parameter, but with the restriction that the type of the parameter must be declared unless it is a number. In technical terms, a setting parameter is translated into an actual variable (as opposed to a definition parameter) which may be dynamically updated.

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Index

Concentric components, 48 Data files, 10 DFT, 57 DFT, 57 Diffraction, 47, 48, 55 Environment variable BROWSER, 11 MCXTAS, 82 MCXTRACE, 11, 75 Keyword Sinculude, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mextrace-r, see Library/Run-time read-table-lib (Read-Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 PROP_DL, 75 Optics, 28, 33, 37	Bugs, 11	PROP_Z0, 75 RESTORE_XRAY, 76	
Data files, 10 DFT, 57 DFT, 57 DFT, 57 Diffraction, 47, 48, 55 Environment variable BROWSER, 11 MCXTAS, 82 MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read-table-lib (Read-Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW-BACKPROP, 75 MC-GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POSA_COMP, 76 POS	Concentric components, 48	ROT_A_COMP, 76	
DFT, 57 Diffraction, 47, 48, 55 SCATTERED, 76 STORE_XRAY, 76 run-time PROP_X0, 75 PROP_Y0, 75 PROP_Y0, 75 Shared, see Library/Components/share Monitor Monitor 4pi, 61 photon counter, 59 Monitors, 58 Banana shape, 68 Custom monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive detector (PSD), coherent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	Data files, 10		
Diffraction, 47, 48, 55 Environment variable BROWSER, 11 MCXTAS, 82 MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 misc, 72 monitors, 58 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CU			
Environment variable BROWSER, 11 MCXTAS, 82 MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read.table-lib (Read.Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW.BACKPROP, 75 MC.GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 7		•	
BROWSER, 11 MCXTAS, 82 MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read.table-lib (Read.Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Monitor 4pi, 61 photon counter, 59 Monitors, 58 Custom monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
BROWSER, 11 MCXTAS, 82 MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read-table-lib (Read-Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW-BACKPROP, 75 MCGETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Monitor Monitor 4pi, 61 photon counter, 59 Monitors, 58 Banana shape, 68 Custom monitoring (user variables, Monitor_nD), 69 Energy monitors, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	Environment variable	PROP_X0, 75	
MCXTAS, 82 MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mextrace-r, see Library/Run-time read-table-lib (Read-Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Monitor 4pi, 61 photon counter, 59 Monitors, 58 Custom monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor nD, 70 Position sensitive detector (PSD), 60 Position sensitive detector (PSD), co-herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	BROWSER, 11		
MCXTRACE, 11, 75 Incoherent elastic scattering, 55 Keyword Keyword Kinclude, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MONITOR, 59 Monitors, 58 Custom monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive detector (PSD), co-herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
$ \begin{array}{llllllllllllllllllllllllllllllllllll$	MCXTRACE, 11, 75	, , , , , , , , , , , , , , , , , , , ,	
Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 misc, 72 monitors, 58 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 MCDISPLAY, 76 GROUP, 76 Monitors, 58 Banana shape, 68 Custom monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), co-herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55		Monitor	
Keyword %include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 monitors, 58 Optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 Monitors, 58 Banana shape, 68 Custom monitoring (user variables, Monitor_nD), 69 Penergy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), co-herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 69 Monitor_nD, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	Incoherent elastic scattering, 55	± ,	
%include, 75 EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 Energy monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive detector (PSD), co-herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 69 Monitor_nD, 69 Energy monitor, 59 Position sensitive detector (PSD), 60 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 69 Monitor_nD, 69 Energy monitor, 59 Position sensitive detector (PSD), 60 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor, 59 monitors PreMonitor_nD, 60 Monitor_nD, 60 Position sensitive detector (PSD), 60 Position sens	T7 1		
EXTEND, 11, 18, 70, 76 GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 Custom monitoring (user variables, Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
GROUP, 76 MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 Monitor_nD), 69 Energy monitor, 59 Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
MCDISPLAY, 76 OUTPUT PARAMETERS, 70 SHARE, 75 Library, 75 Components data, 10, 82 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 Monitor, 59 Photon parameter correlations, Pre-Monitor, D Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), co-herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
OUTPUT PARAMETERS, 70 SHARE, 75 Components data, 10, 82 monitors, 58 copties, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 SHORD ARC MERCE Photon parameter correlations, Pre-Monitor_nD, 70 Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), coherent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55		· · · · · · · · · · · · · · · · · · ·	
SHARE, 75 Library, 75 Components data, 10, 82 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 Library, 75 Position sensitive detector (PSD), coherent, 61 Position sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55		= -	
Library, 75 Components data, 10, 82 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 Category Position sensitive detector (PSD), 60 Position sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 60 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
Library, 75 Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Camponents Position sensitive detector (PSD), coherent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	SHARE, 75		
Components data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 herent, 61 Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	Library 75	· / /	
data, 10, 82 misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 monitors Position Sensitive Detector (PSD), Energy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55		· /·	
misc, 72 monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 mensitive between (15D), En ergy Selective, 62 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55 Multiple scattering, 55	•		
monitors, 58 optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Position sensitive monitor recording mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
optics, 28 samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 mean energy, 69 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
samples, 43 share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 sources, 18 PowerPSD, 62 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	•		
share, 75 sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 The All-in-One monitor (Monitor_nD), 64 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55	- · · · · · · · · · · · · · · · · · · ·		
sources, 18 mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 monitors Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
mcxtrace-r, see Library/Run-time read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 Wavelength monitor, 59 monitors PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
read_table-lib (Read_Table), 10, 80 Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Run-time, 75 PreMonitor, 70 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
Run-time, 75 ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Run-time, 75 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55		· ·	
ABSORB, 75 ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 Monte Carlo method, 12 Accuracy, 17 Direction focusing, 15 Stratified sampling, 16 Multiple scattering, 55			
ALLOW_BACKPROP, 75 MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 Multiple scattering, 55			
MC_GETPAR, 70, 76 NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 POS_A_CURRENT_COMP, 76 Multiple scattering, 55			
NAME_CURRENT_COMP, 76 POS_A_COMP, 76 POS_A_CURRENT_COMP, 76 Multiple scattering, 55 POS_A_CURRENT_COMP, 76		• /	
POS_A_COMP, 76 Multiple scattering, 55 POS_A_CURRENT_COMP, 76			
POS_A_CURRENT_COMP, 76	•	- 9.	
PROP_DL, 75 Optics, 28 , 33 , 37			
	PROP_DL, 75	Optics, 28, 33, 37	

```
Beam stop, 29
    Filter, 30
    lens, 31, 33-36
    mirror, 37, 38
    Mirror plane, 38
    multilayer, mirror, 38
    Point in space (Arm, Optical bench, Co-
         ordinate system), 28
    Progress bar, 73
    Slit, 28
Removed x-ray events, 76
Removed xray events, 10
Sample environments, 48
Samples, 43
    Absorption, 45
    Dilute colloid medium, 47
    liquid, diffraction, 57
    Powder, multiple diffraction line, 48
    SAXS, 47
    Single crystal diffraction, 55
    single crystal, darwin, 54
Small angle scattering, 47
Sources, 18
    Continuous source with specified diver-
         gence, 21
    Flat surface source, 21
    Point source, 20
    Shadow, 74
    Source_gaussian, 22
    Virtual source, 74
    Virtual source from stored neutron
         events, 73
    Virtual source, recording photon events,
    X-ray tube laboratory source, 23
time resolved, 57
Tomography, 45
Tools
```

mcdoc, 11